The HERMES project

High Energy Rapid Modular Ensamble of Satellites

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The Gamma-Ray Burst phenomenon

- sudden and unpredictable bursts of hard-X / soft gamma rays with huge flux
- most of the flux detected from 10–20 keV up to 1–2 MeV,
- fluences for very bright GRB (about 3/yr) 25 counts/cm²/s (GRB 130427A 160 counts/cm²/s)
- bimodal distribution of duration (0.1–1.0 s & 10.0–100.0 s)
- measured rate (by an all-sky experiment on a LEO satellite): ~ 0.8 /day (estimated true rate ~ 2 / day)
- evidence of submillisecond structures



The Gamma-Ray Burst phenomenon

Prompt Emission:

Short: $\tau \approx 0.2$ sec, Fluence $\approx 4 \times 10^{-7} \text{ erg/cm}^2 (25 \text{ keV} - 1 \text{MeV})$

=> Binary NS mergers (GW sources)

Long: $\tau \approx 25$ sec, Fluence $\approx 8 \times 10^{-6} \text{ erg/cm}^2 (25 \text{ keV} - 1 \text{MeV})$

=> Hypernovae (SNe Massive Stars)







Counts / 31.25 ms

80

40

The Gamma-Ray Burst phenomenon

Millisecond variability (minimum variability time-scale, MacLachlan et al. 2013) Short: 3 msec (wavelet techniques)

Long: 30 msec (wavelet techniques)

Internal shock model (ultarelativistic, $\gamma \approx 10^2 \div 10^3$, colliding shocks)

GRB Luminosity Function: <8 c/s (50-300) keV short

> 8 c/s (50–300) keV long





Number of GRB and Fluxes

Short GRBs:

Duration: 0.2 sec,

Counts (50-300 MeV): 8 c/cm²/s

Averaged photon energy: $(Emax x Emin)^{1/2} = 122 \text{ keV}$

Fluence: $0.2 \ge 8 \ge 122 \text{ keV/cm}^2 = 3 \ge 10^{-7} \text{ erg/cm}^2$

Fermi GBM - 4-years data

14 Short GRB burst per year with

count rate > 8 c/s





Simulations of a bright short GRB (50 - 300 keV)Background: 0.43 c/s/cm²/steradiansBackground for 2 steradians FOV: 0.86 c/cm²/sProton fluxes in LEO (580 km): 0.165 c/cm³/sActivation in equatorial LEO (580 km): $\leq 0.3 \text{ c/cm}^3$ /s (not included)Burst duration: 0.2 secSource count rate: 7.875 ph/cm²/sExponential shot rate: 100 shot/sBand 50-300 keVExponential shot decay time: 1 msec





Delays from cross-correlation analysis

Cross-correlation of GRB lightcurves from two satellites of 100 cm2 effective area in the 50-300 keV band:



Simulation	Radius [cm]	Expected delay [s]	Measured delay [s]	Error [s]	Error in unit of σ
sim_1.fits	1·10 ⁹	$6.67128190 \cdot 10^{-2}$	$6.66744709 \cdot 10^{-2}$	$0.0561 \cdot 10^{-3}$	0.6830805064
sim_2.fits	$9.9 \cdot 10^8$	$6.60456908 \cdot 10^{-2}$	$6.59104213 \cdot 10^{-2}$	$0.1233 \cdot 10^{-3}$	1.0970766361
sim_3.fits	9·10 ⁸	$6.00415371 \cdot 10^{-2}$	$5.98946773 \cdot 10^{-2}$	$0.1759 \cdot 10^{-3}$	0.8349052625
sim_4.fits	$8.1 \cdot 10^8$	$5.40373834 \cdot 10^{-2}$	$5.40064611 \cdot 10^{-2}$	$0.0791 \cdot 10^{-3}$	0.3911742201
sim_5.fits	$7.2 \cdot 10^8$	$4.80332297 \cdot 10^{-2}$	$4.80094887{\cdot}10^{-2}$	$0.0735 \cdot 10^{-3}$	0.3228751331
sim_6.fits	$6.3 \cdot 10^8$	$4.20290760 \cdot 10^{-2}$	$4.19750251 \cdot 10^{-2}$	$0.0662 \cdot 10^{-3}$	0.8163554594
sim_7.fits	$5.4 \cdot 10^8$	$3.60249223 \cdot 10^{-2}$	$3.59473675 \cdot 10^{-2}$	$0.0646 \cdot 10^{-3}$	1.2007242824
sim_8.fits	$4.5 \cdot 10^8$	$3.00207686 \cdot 10^{-2}$	$2.9960330 \cdot 10^{-2}$	$0.0764 \cdot 10^{-3}$	0.7907702189
sim_9.fits	3.6·10 ⁸	$2.40166149 \cdot 10^{-2}$	$2.39778887 \cdot 10^{-2}$	$0.0709 \cdot 10^{-3}$	0.5464393152
sim_10.fits	$2.7 \cdot 10^8$	$1.80124611 \cdot 10^{-2}$	$1.79572739 \cdot 10^{-2}$	$0.0612 \cdot 10^{-3}$	0.9024896271
sim_11.fits	$1.8 \cdot 10^8$	$1.20083074 \cdot 10^{-2}$	$1.19955540 \cdot 10^{-2}$	$0.0864 \cdot 10^{-3}$	0.1475920279
sim_12.fits	9·10 ⁷	$0.60101537 \cdot 10^{-2}$	5.99619420.10-3	$0.0766 \cdot 10^{-3}$	0.1822152926

Error in cross-correlation accuracy: 84 μ sec Number of independent estimate of delays: Nsatellite – 1 Position of the source in the sky, (α , δ): 2 parameters Statistical improvement in determining the position in the sky with Nsatellite: (Nsatellite –1– 2)^{1/2} = 8.5 Error in delay accuracy: 8.5 μ sec (Nsatellite = 100) 12 μ sec (Nsatellite = 50)



Determination of source position through delays

Error in accuracy \approx c \times (error in delay accuracy / average baseline)

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Maximum baseline = 2 \times (\text{Rearth} + \text{Hsatellite}) = 2 \times (6371 + 580) \text{ km}
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Average baseline = Maximum baseline / 2

Error in accuracy = 75 arcsec (for Nsatellite 100)

Error in accuracy = 110 arcsec (for Nsatellite 50)



Detector and satellite

Detector

Scintillator Crystals: CsI (classic) or LaBr₃ or CeBr₃ (rise – decay: 0.5 - 20 ns) Photo-detector: Silicon Photo Multiplier (SiPM) or Silicon Drift Detector (SDD) Effective area: 10×10 cm Crystal thickness: 1 cm Weight: 0.5 - 1 kgEnergy band: 50 - 300 keVEnergy resolution: 15% at 30 keV Temporal resolution: ≤ 10 nanoseconds **Satellite** 5 detectors on a cubic structure + solar panel Weight: $\leq 10 \text{ kg}$ Shielding Grating shields to reduce proton flux to $0.165 \text{ c/cm}^3/\text{s}$ **Collimator** 2 stearadians (0.6 stearadians Icosahedron 20 faces, 0.13 stearadians Snub Dodecahedron 92 faces, strong reduction of X-ray background) **Data recording** Continuous recording of buffered data





The HERMES mission

High Energy Rapid Modular Experiment Satellites (a nanosatellite swarm monitor for GRB & High Energy GW counterparts) **GRB** statistics Average GRBs: 300/yr

Bright GRBs: 30/vr GRB structure: duration 25 s, shot noise $\tau = 1$ ms, rate = 100/s Instrument $N \ge 100$ Nano Satellites (Modules) in Low Earth Orbit Average separation between Modules: 6000 km Module (weight ≤ 10 kg)

5 Detectors

Field of View of each Detector: 2 steradians GPS absolute temporal accuracy < 100 nanoseconds GPS based Module positional accuracy: $\leq 10 \text{ m}$

Detector

Scintillator Crystals: CsI (classic) or LaBr₃ or CeBr₃ (rise – decay: 0.5 - 20 ns) Photo-detector: Silicon Photo Multiplier (SiPM) or Silicon Drift Detector (SDD) Effective area: 10×10 cm Weight: 0.5/1 kg Energy band: 3 keV – 50 MeV Energy resolution: 15% at 30 keV Temporal resolution: ≤ 10 nanoseconds **Mission performance** Accuracy in delays between Average GRB lightcurves of two Modules (cross correlation techniques): 20 microseconds for Average GRBs Continuous recording of buffered data Triggered to ground telemetry transmission Accuracy in positioning of Bright GRBs: 75 – 110 arcsec Range of accuracy in positioning of GRB: from 25 to 330 arcsec Modular structure: overall effective area 1 m² every 100 modules



Scintillator Crystal

The Uncertainty Relation $\Delta r \Delta t > G\hbar/c^4$ and the space-time diagram for the intervals (Burderi, Di Salvo, Iaria, Physical Review D, 93, 064017, 2016)



The new Uncertainty Principle and the Minkowski metric: preserving Lorentz Invariance



GRB & Quantum Gravity (Massive Photons or Lorentz Invariance Violation)

MP or LIV predictions:

 $|v_{phot}/c - 1| \approx \xi E_{phot}/(M_{QG} c^2)^n$ ($\xi \approx 1$ n = 1,2) and $M_{QG} = \zeta m_{PLANCK}$ ($\zeta \approx 1$)

$$\Delta t_{\text{MP/LIV}} = \xi \left(D_{\text{TRAV}}/c \right) \left[\Delta E_{\text{phot}}/(M_{\text{QG}} c^2) \right]^n$$
$$D_{\text{TRAV}}(z) = (c/H_0) \int_0^z d\beta (1+\beta) / \left[\Omega_\Lambda + (1+\beta)^3 \Omega_M \right]^{1/2}$$

Band		Flux	Fluence	Expected Δt_{OG}	Expected $\Delta t_{OGR} \propto D_{GRB}/c$	
		(Bright GRBs) $(1 \text{ m}^2, 10 \text{ s})$		for Quantum Grav	for Quantum Gravity effects	
				z = 0.9	z = 3.0	
(keV)		$(counts/cm^2/s)$	(counts)	(μs)	(µs)	
2 -	25	24.7	2,470,000	0	0	
25 -	50	6.2	620,000	1	2	
50 -	100	5.5	550,000	2	3	
100 -	300	6.1	610,000	3	5	
300 -	1000	2.4	240,000	12	19	
1000 -	2000	0.4	40,000	28	45	
2000 -	5000	0.15	15,000	65	104	
5000 -	50000	0.07	7,000	421	671	

Conclusions I

All sky monitor of Gamma Bursts (GRB, Magnetar, High Energy counterparts of Gravitational Waves, etc.)

Accuracy in positioning of Bright GRBs: 75 – 110 arcsec

range of accuracy in positioning of GRB: from 25 to 330 arcsec

 1 m^2 effective area (50 – 300 keV)

Energy resolution: 15% at 30 keV

Temporal resolution: ≤ 10 nanoseconds

Quantum Gravity: probing the ultimate structure of space-time

Time lags

caused by prompt emission mechanism: complex dependence from E_{phot} (Band II) and E_{phot} (Band I)

independent of $D_{GRB}(z_{GRB})$

caused by Quantum Gravity effects:

 $\propto |E_{phot}(Band II) - E_{phot}(Band I)|$

 $\propto D_{GRB}(z_{GRB})$

the two effects can be disentangled with experimentally measured:

 Δt_{OGR} (HERMES)

 z_{GRB} (optical, follow-up observations of host galaxy)

Conclusions II

1) Cheap:

simple detector & nano(small)satellites: up to 100 million € for 100 satellites see e.g. Thales Alenia Space: 40 kg - 100 W, 3 axes pointing, LEO, cost ≈ 1 M€ ("deep throat", private comm.)

2) Fast:

few years (\leq 5 years) to flight the first satellite(s)

3) Modular:

robust against one or more satellite(s) failure

Growing interest in constellation of small satellites...



August 22 - 26, 2016, Innsbruck, Austria

2nd BRITE-Constellation Science Conference "small satellites - big science"

That's all Folks!