

WE LOOK AFTER THE EARTH BEAT



METIS DESIGN OVERVIEW AND TECHNICAL CHALLENGES

**METIS: 2nd Science and Technical Meeting
Turin, 12-13 December 2012**

13/12/2012



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METIS MISSION CONTEXT: SOLAR ORBITER

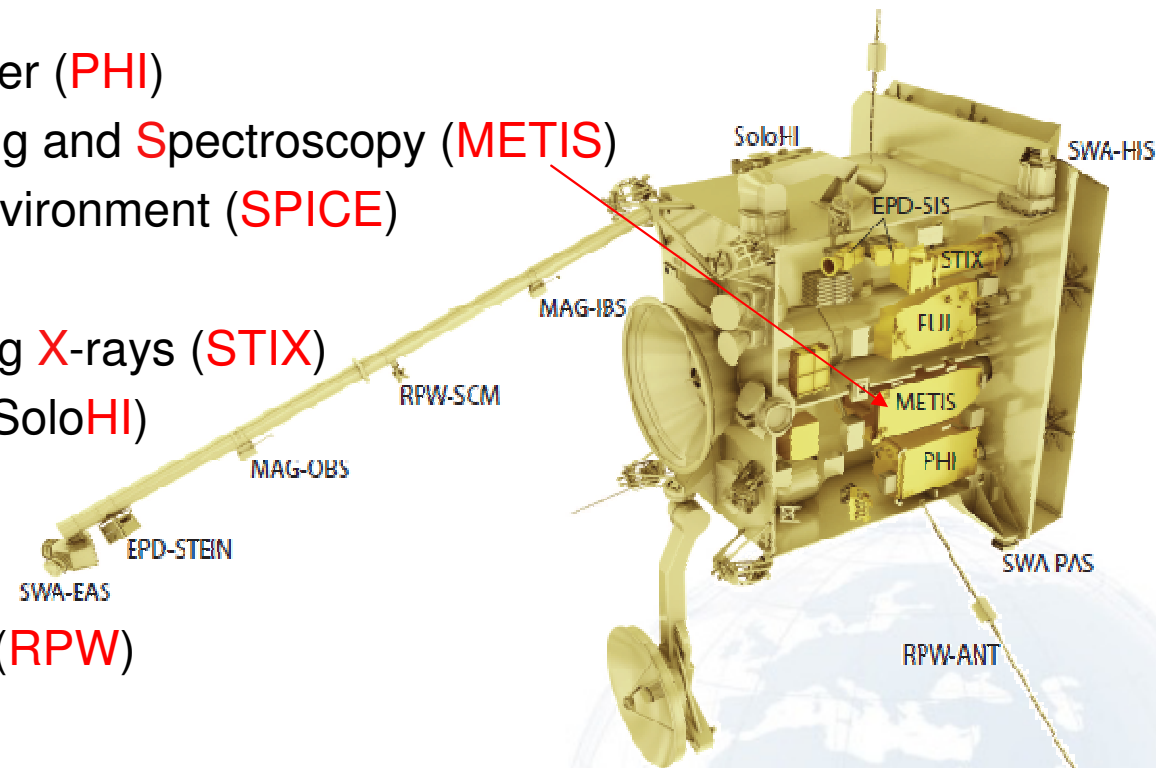
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- First mission of ESA's Cosmic Vision 2015–2025 programme
- Objectives: exploration of the Sun and of the inner heliosphere combination of in-situ and remote-sensing instruments.
- Remote-sensing instruments:

1. Polarimetric and Helioseismic Imager (PHI)
2. Multi-Element Telescope for Imaging and Spectroscopy (METIS)
3. SPectral Imaging of the Coronal Environment (SPICE)
4. Extreme Ultraviolet Imager (EUI)
5. Spectrometer/Telescope for Imaging X-rays (STIX)
6. Solar Orbiter Heliospheric Imager (SoloHI)

➤ In-situ instruments:

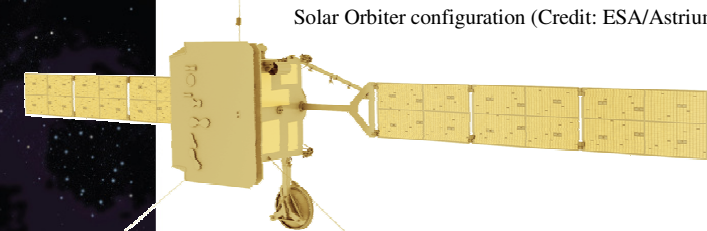
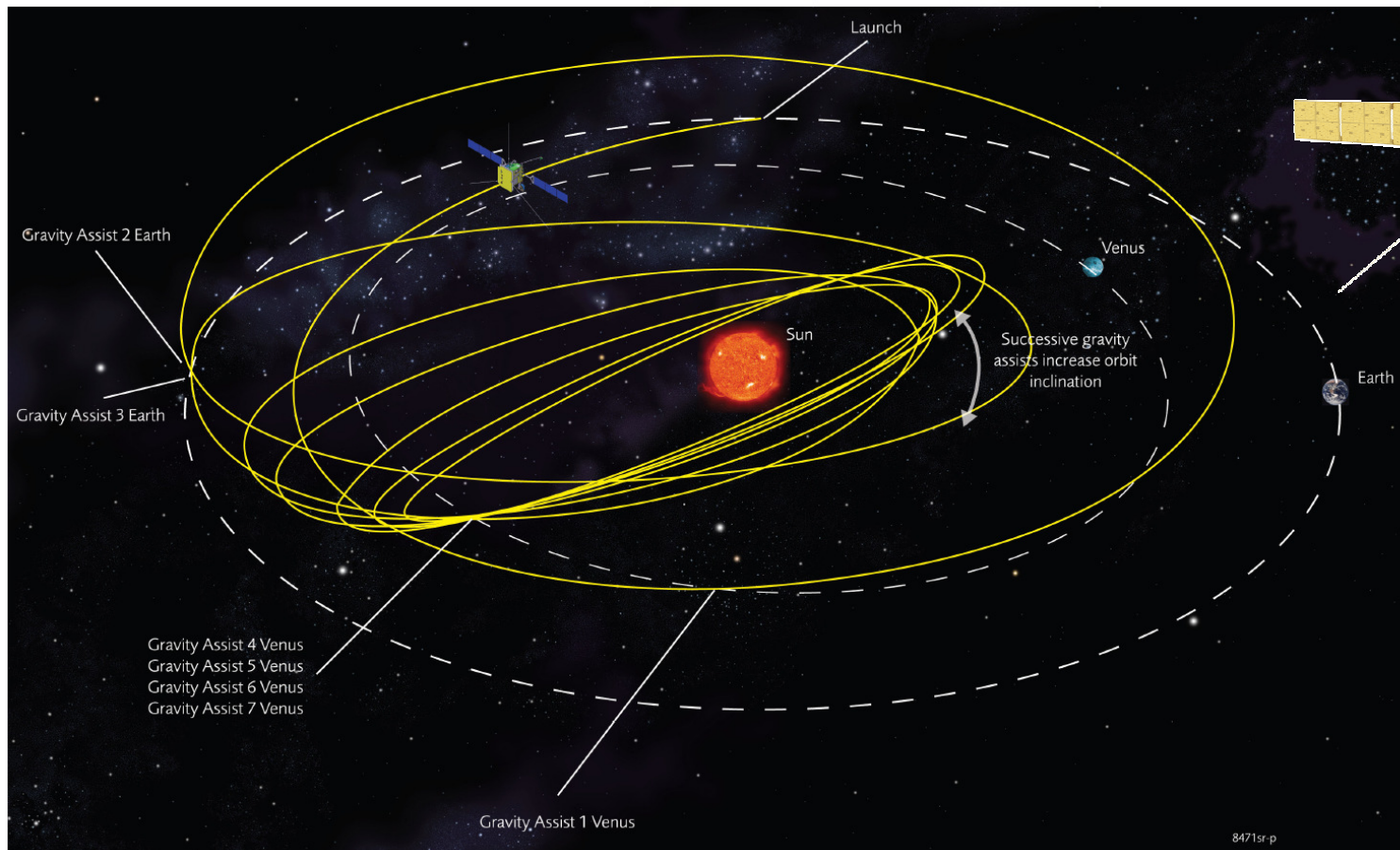
1. Solar Wind Analyser (SWA)
2. Radio & Plasma Wave experiment (RPW)
3. MAGnetometer (MAG)
4. Energetic Particle Detector (EPD)



Solar Orbiter configuration (Credit: ESA/Astrium)

SOLAR ORBITER MISSION PROFILE

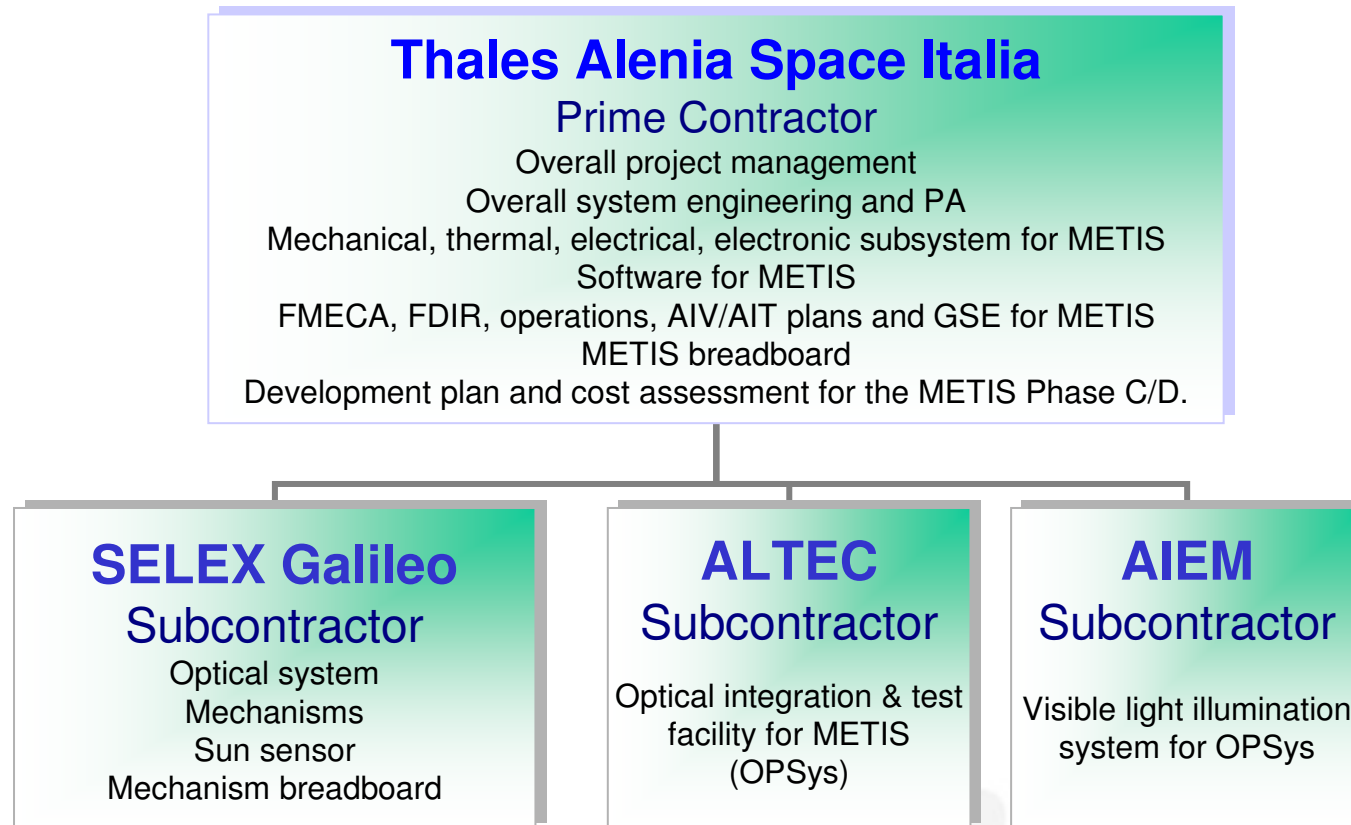
- Planned launch date: January 2017
- Mission duration: 7 years (nominal, including 3.4-year cruise) +3 years (extension)
- Closest perihelion: 0.28 AU, achieved after 4 gravity assist manoeuvres
- Maximum solar latitude: 34°, achieved after other 3 gravity assist manoeuvres



- 3-axis stabilized platform based on reaction wheels and chemical propulsion
- Solar arrays that can be rotated about their longitudinal axis to manage the cell temperature
- Articulated high-gain antenna in X-band
- Heat Shield keeping in shadow the spacecraft body

METIS DEVELOPMENT STATUS AND INDUSTRIAL TEAM

- Phase B industrial contract assigned by Italian Space Agency to a team led by Thales Alenia Space Italia on June 2011
- Preliminary Design Review accomplished on June 2012
- Phase B Final Review established on 20 December 2012



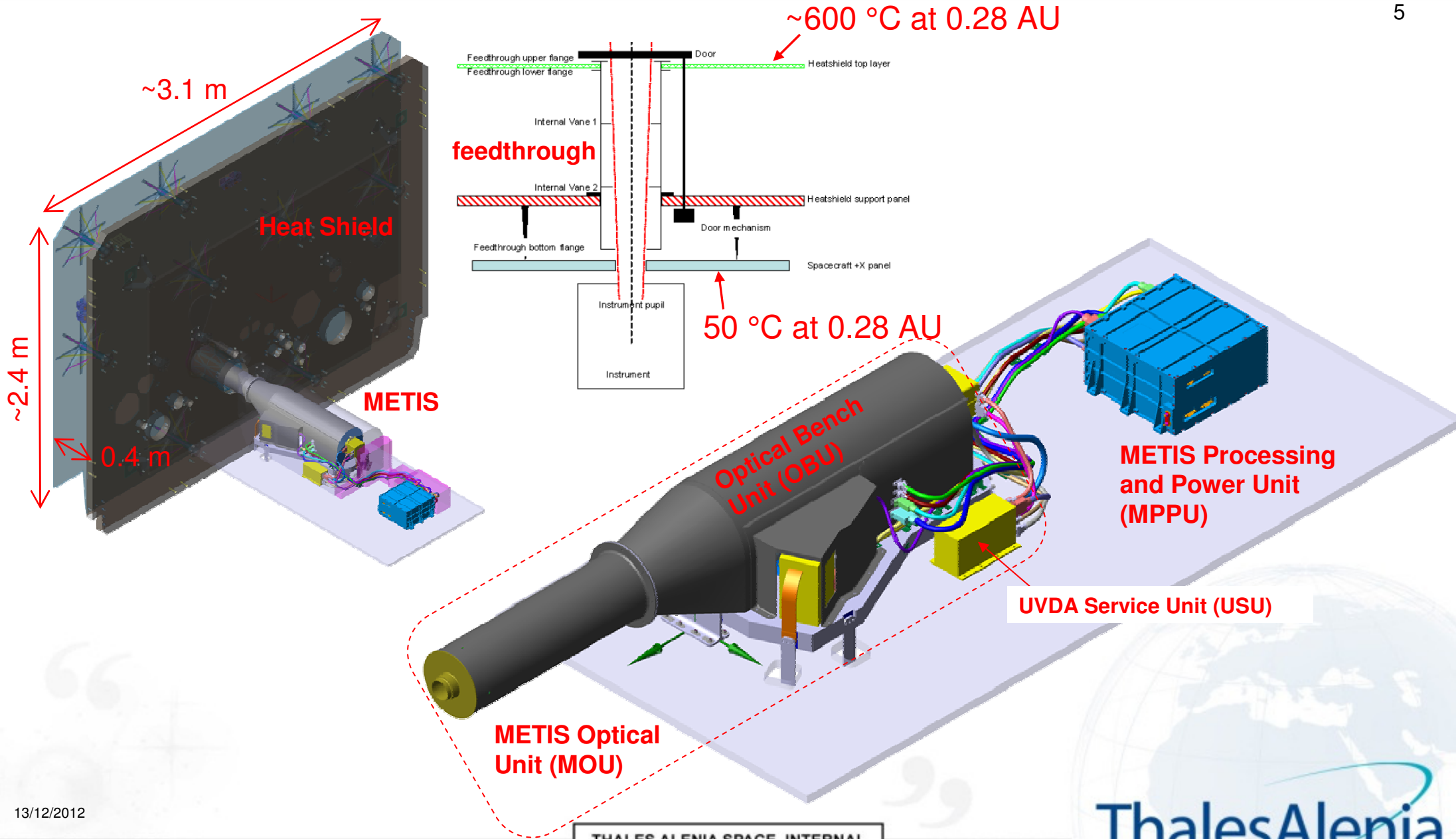
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METIS DESIGN OVERVIEW



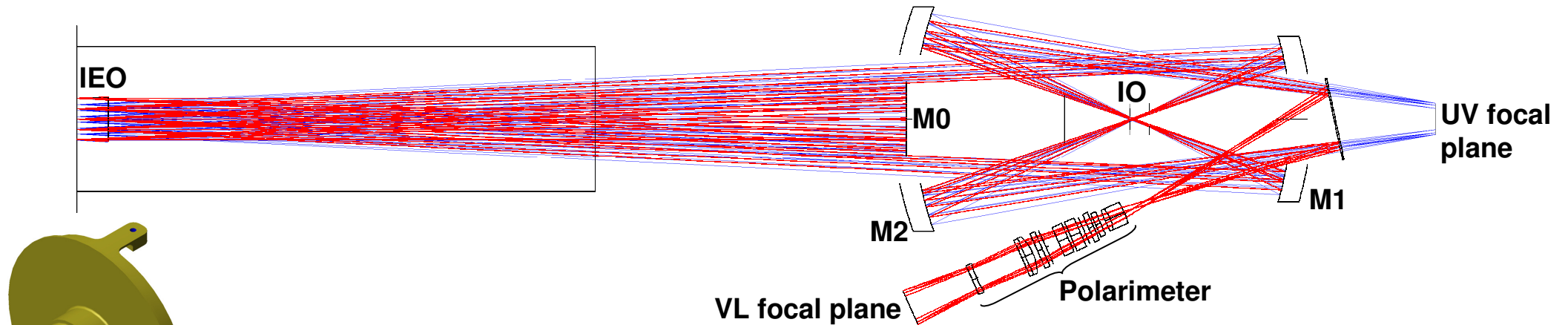
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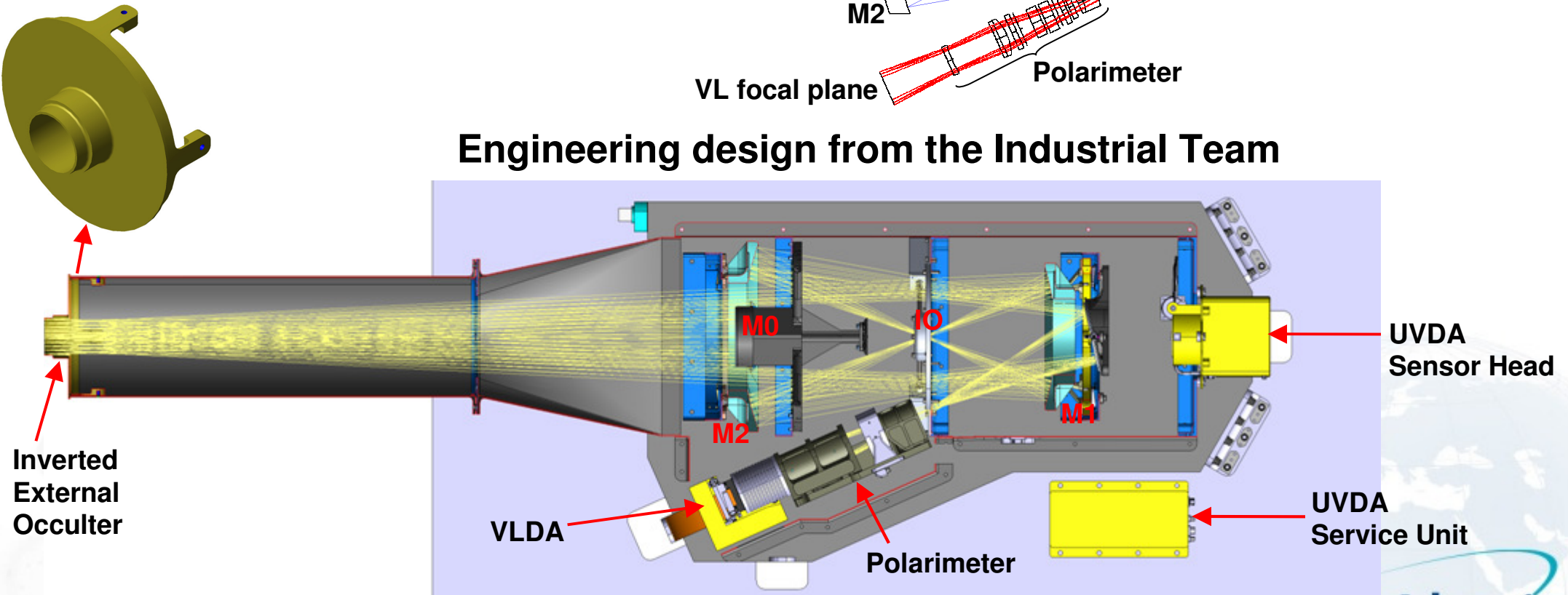
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METIS DESIGN OVERVIEW

Optical design from the Scientific Team



Engineering design from the Industrial Team

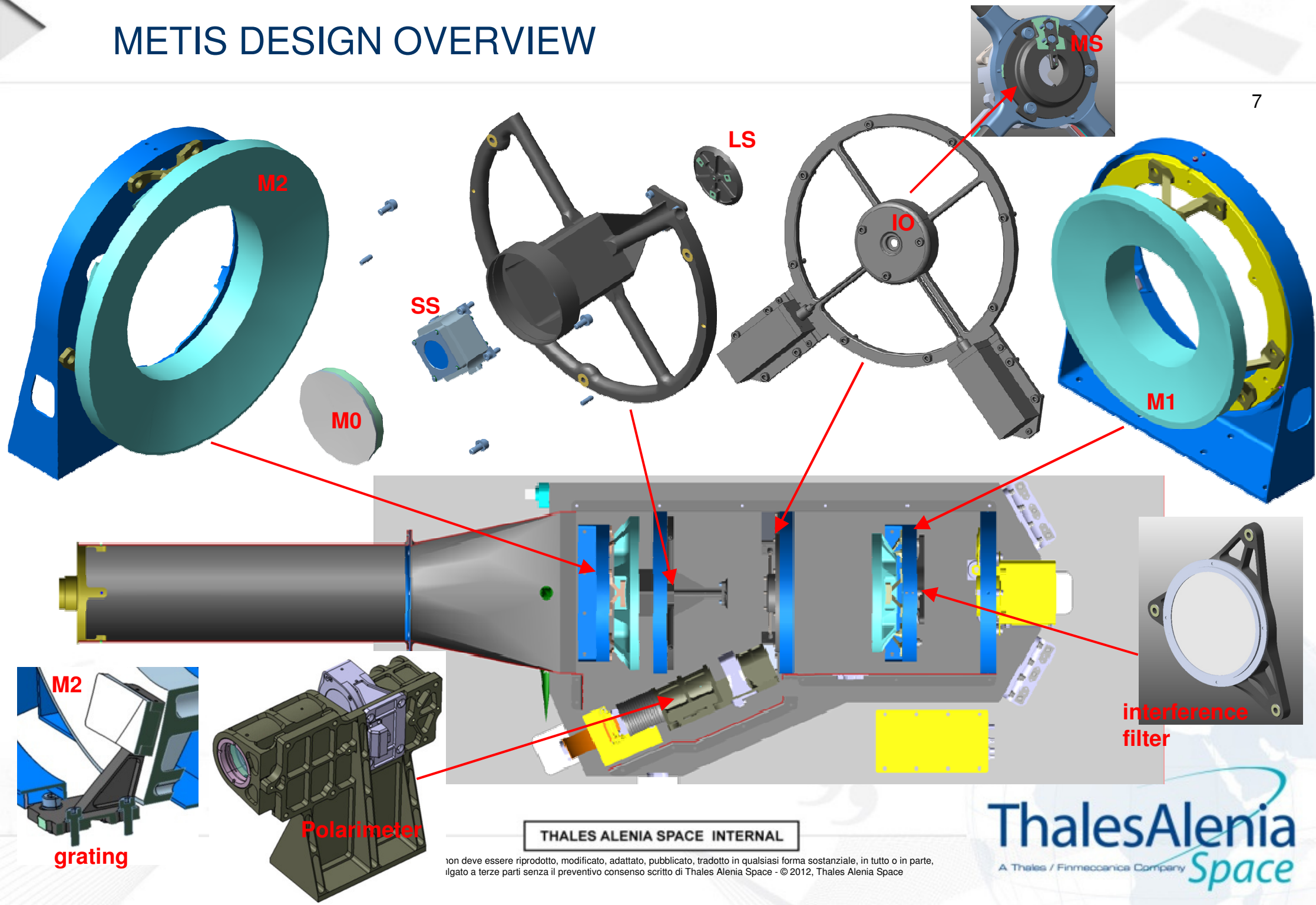


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METIS DESIGN OVERVIEW



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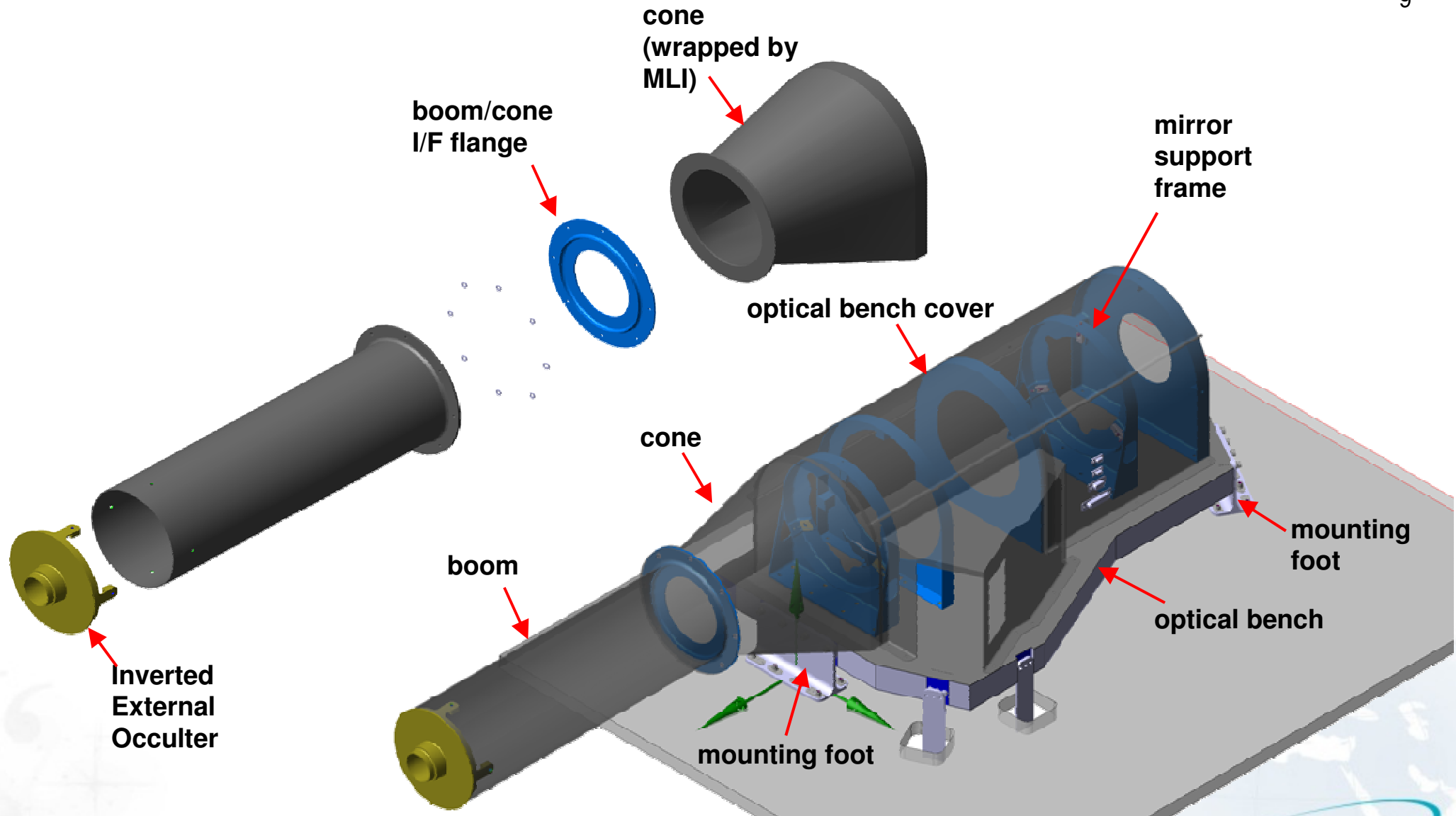
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METIS TECHNICAL CHALLENGES

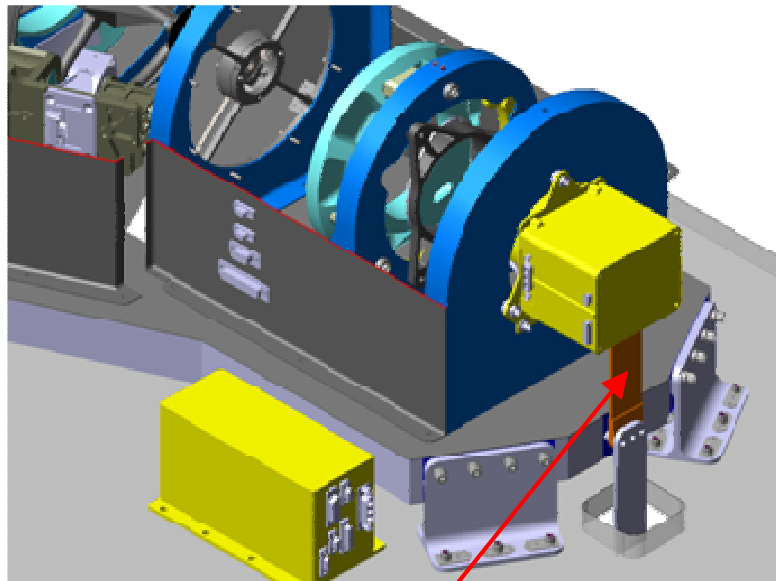
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- Stray-light suppression at the visible detector below 10^{-9} of the Sun disk irradiance (below 10^{-7} at the UV detector). →
 - Very precise shaping of the occulting elements: edge curvature radius $< 50 \mu\text{m}$.
 - Extreme polishing of the mirror surfaces: roughness = 0.3 nm rms (0.2 nm goal).
 - Extreme cleanliness levels at instrument delivery: 3.3 ppm particle cleanliness, 100 ng/cm² molecular cleanliness on mirrors and stops surfaces; 3.3 ppm, 200 ng/cm² on Inverted External Occulter and telescope interior.
 - All instrument internal surfaces (except mirrors) with “black”, non-specular coating/finishing.
- Need to operate the detectors at low temperature (-80 °C for the visible sensor and -20 °C, goal, for the UV sensor) in an extremely severe thermal environment: instrument aperture subject to a thermal flux of ~ 13 solar constants at 0.28 AU from the Sun.
- Dimensional stability $\leq 15 \mu\text{m}$ between mirrors M1, M2 from launch till end of mission lifetime.
- Limited resources (mass, power, allowable conductive and radiative heat rejection) allocated by Solar Orbiter to the METIS instrument.
- All fundamental resonance frequencies of METIS above 140 Hz.
- Capability to withstand a severe random vibration environment: 15.8 g rms \perp , 12.8 g rms //
- Alignment of an optical system without a materialised optical axis (the mirrors M1 and M2 are circular corona) and with UV/EUV channels.

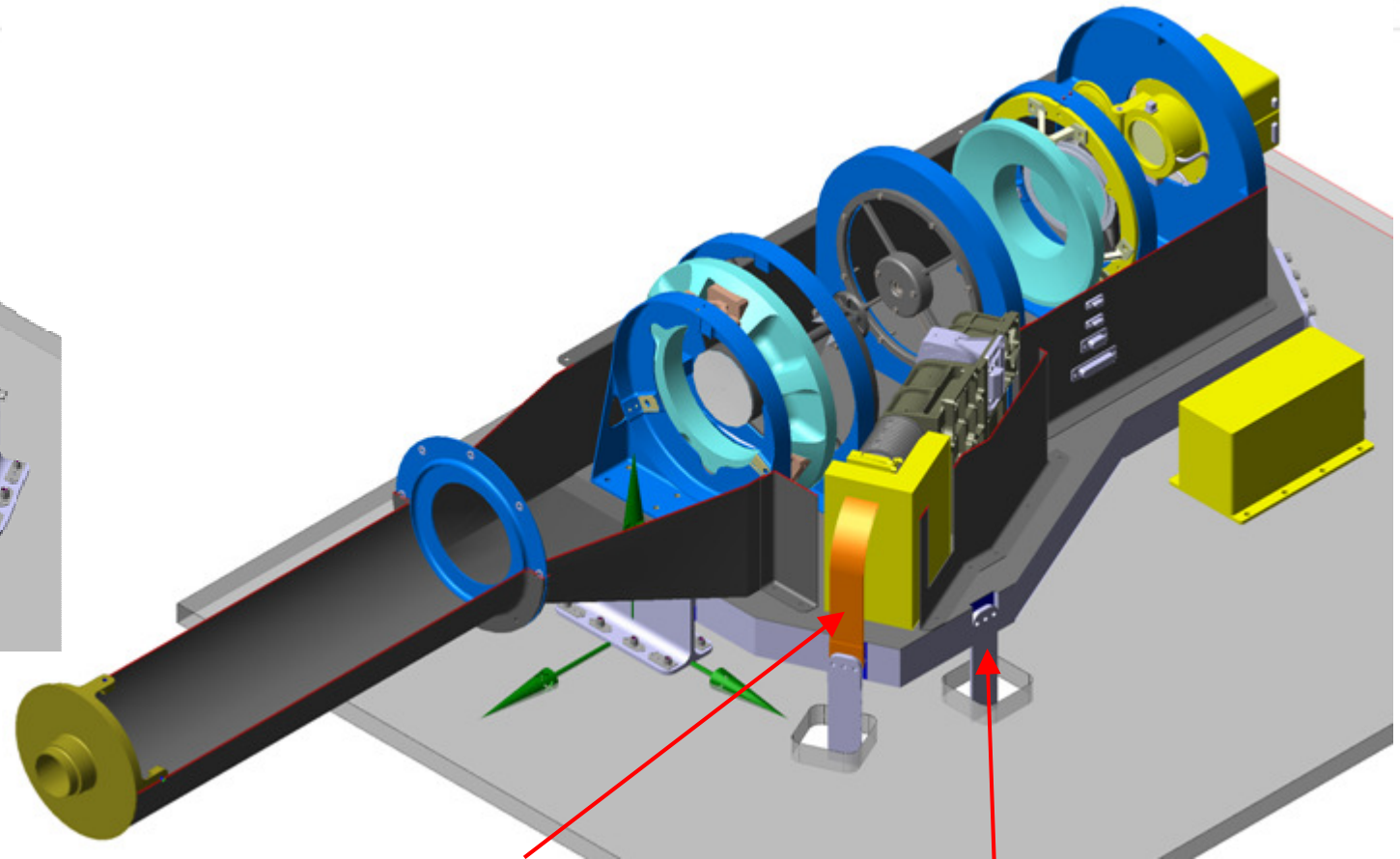
METIS THERMO-MECHANICAL DESIGN



METIS THERMO-MECHANICAL DESIGN



Thermal strap for heat transfer from UV sensor to the S/C radiator for “medium elements” (I/F temperature at 0.28 AU: -10 °C; -30 °C required for cooling the sensor at -20 °C)



Thermal strap for heat transfer from VL sensor TEC to the S/C radiator for “cold elements” (I/F temperature at 0.28 AU: -60 °C)

Thermal strap for heat transfer from the optical bench to the S/C radiator for “hot elements” (I/F temperature at 0.28 AU: 50 °C)

Power consumed by the Thermo-Electric Cooler (TEC) for keeping the VL sensor at -80 °C at 0.28 AU: ~7 W

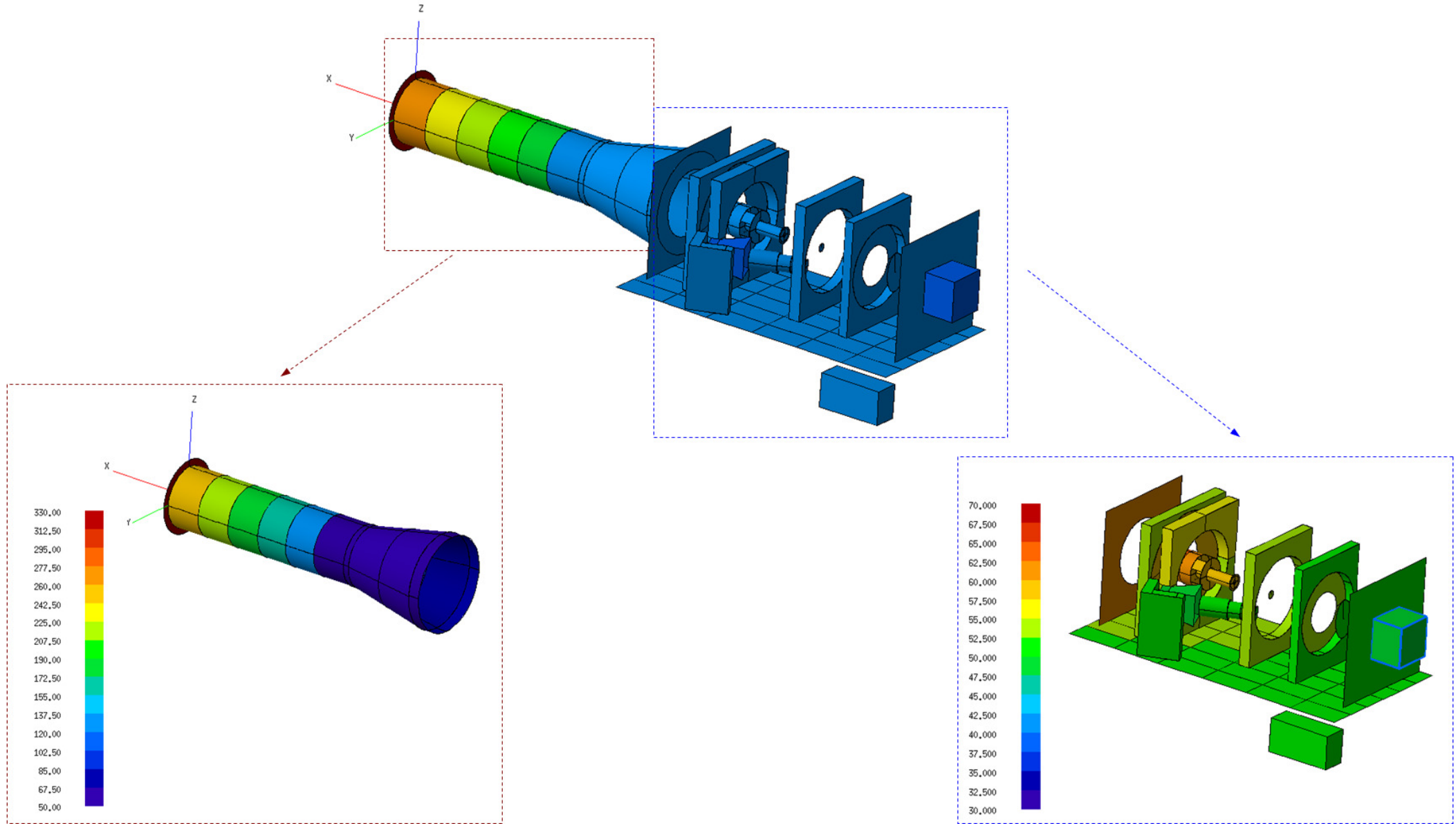
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TEMPERATURE MAP: HOT OPERATIONAL CASE AT 0.28 AU



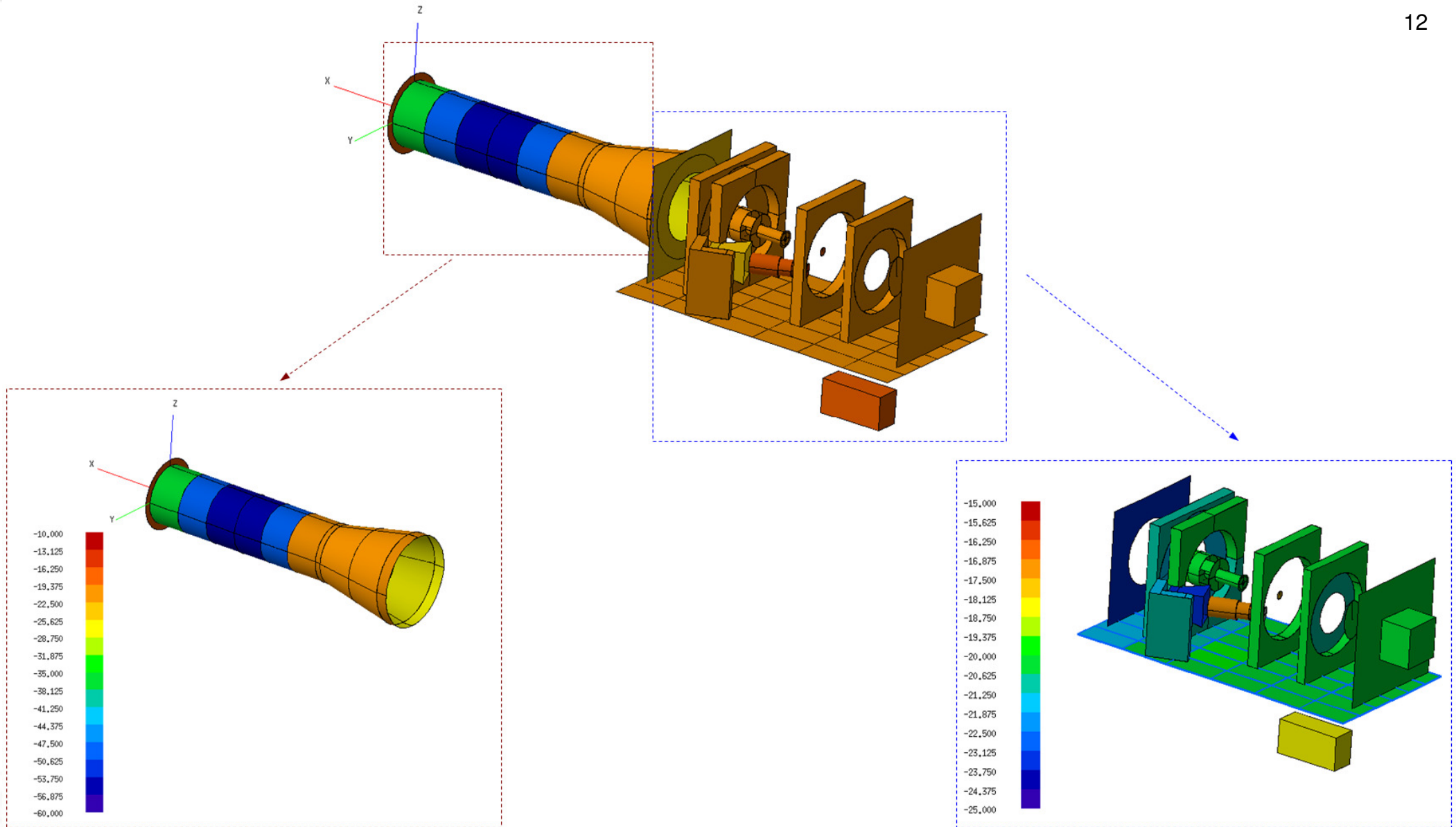
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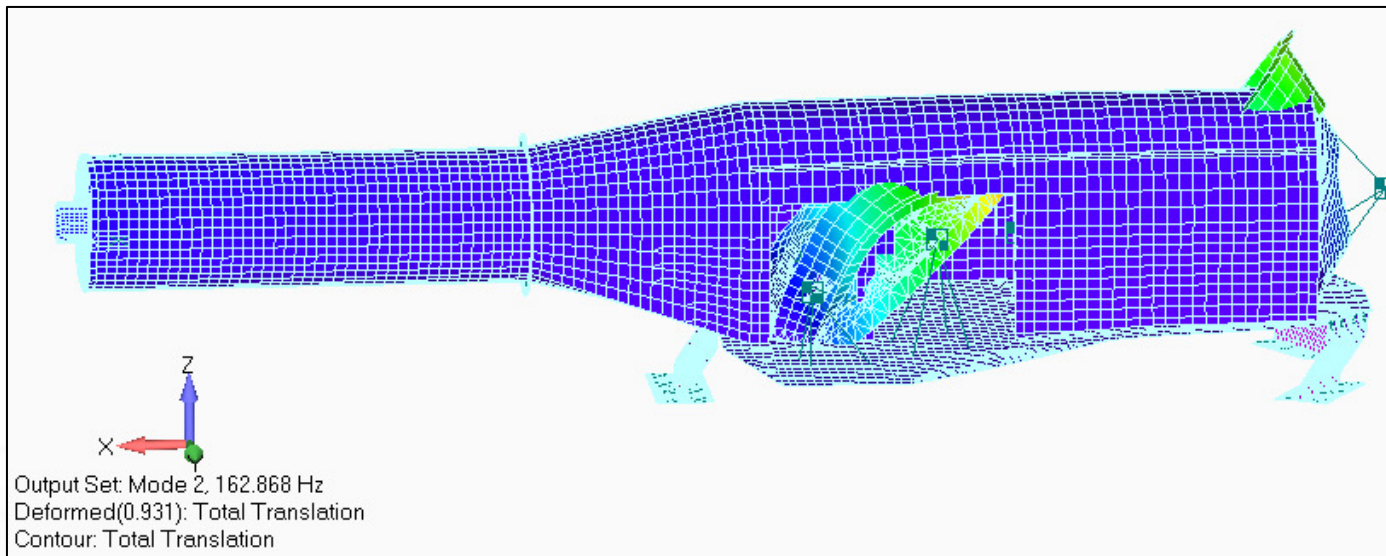
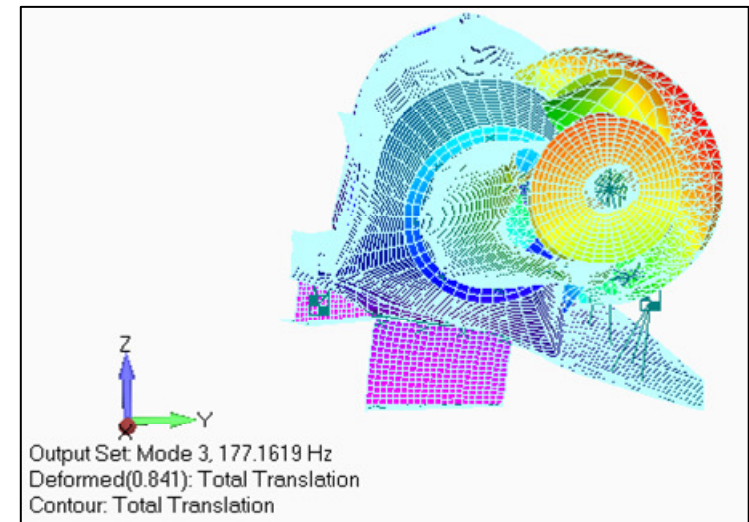
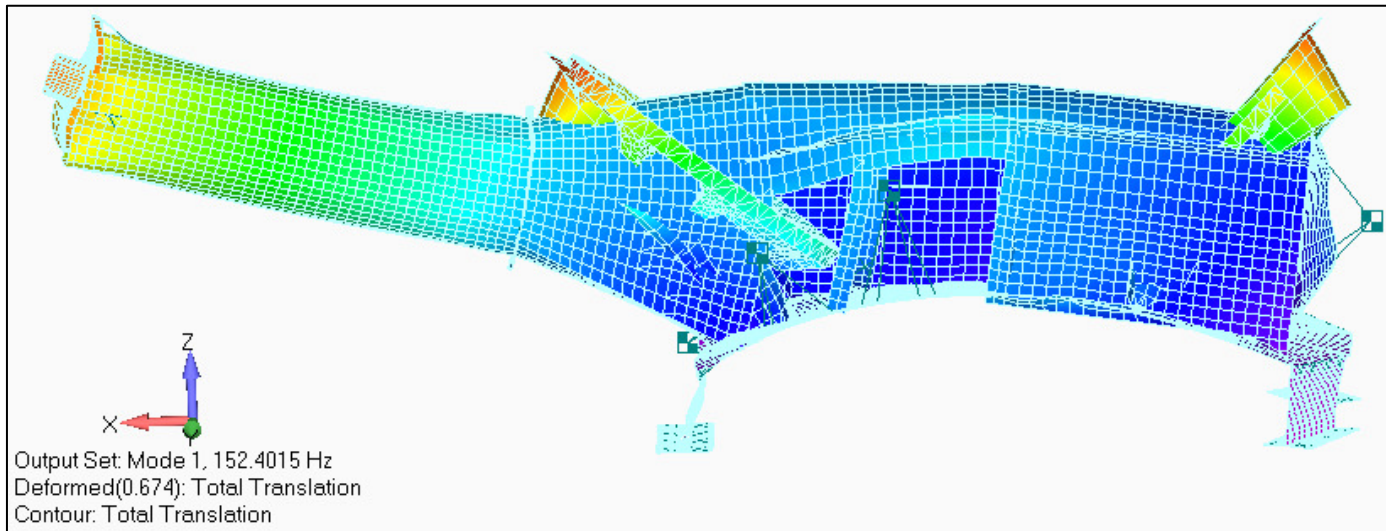
TEMPERATURE MAP: COLD OPERATIONAL CASE AT 0.952 AU



METIS MECHANICAL BEHAVIOUR

First three global vibration modes of METIS

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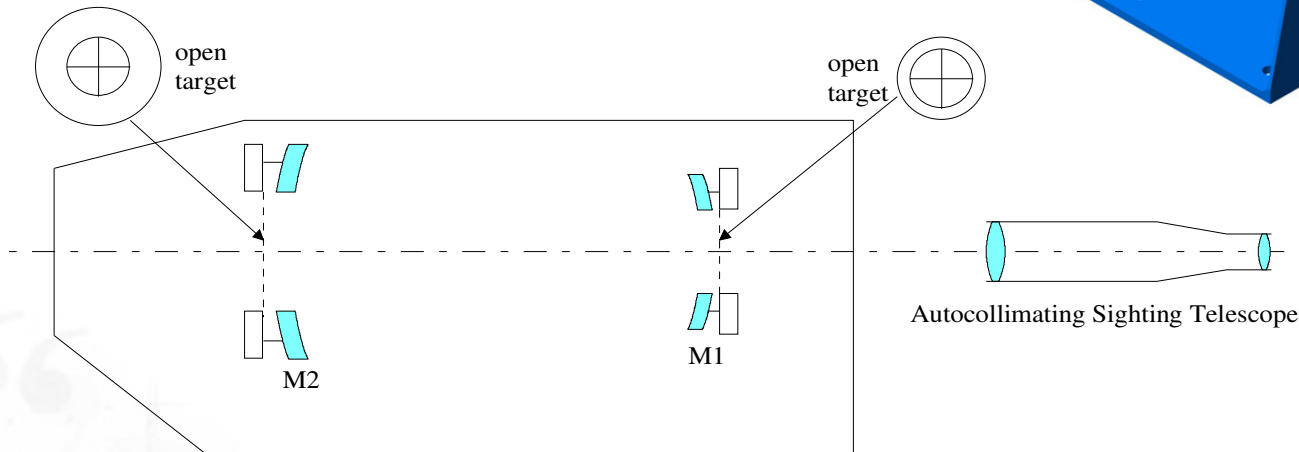
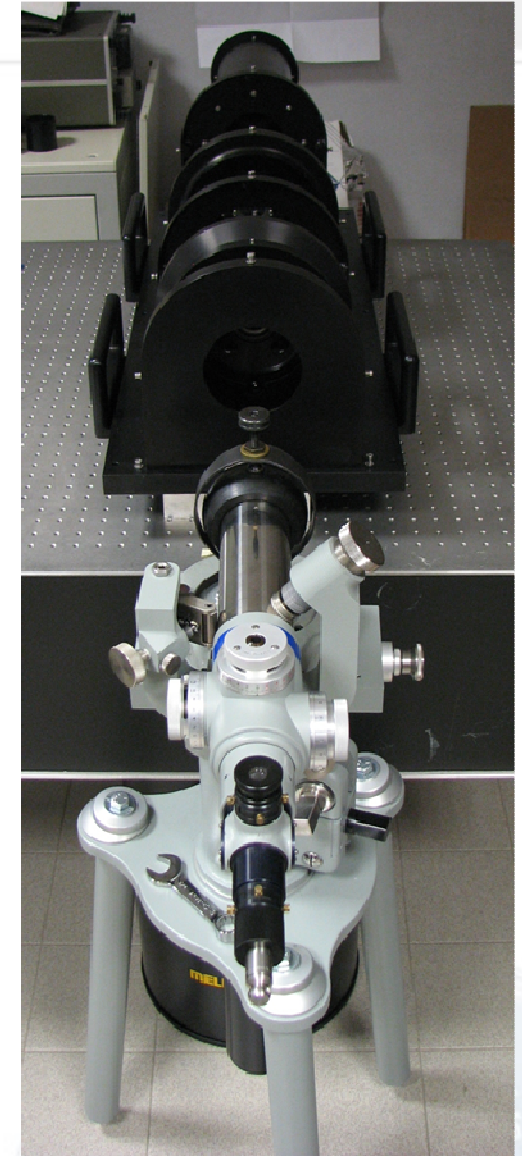
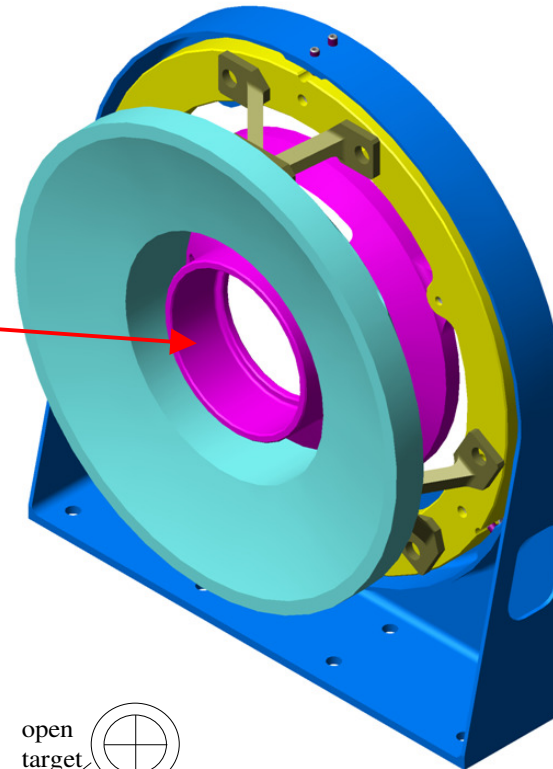
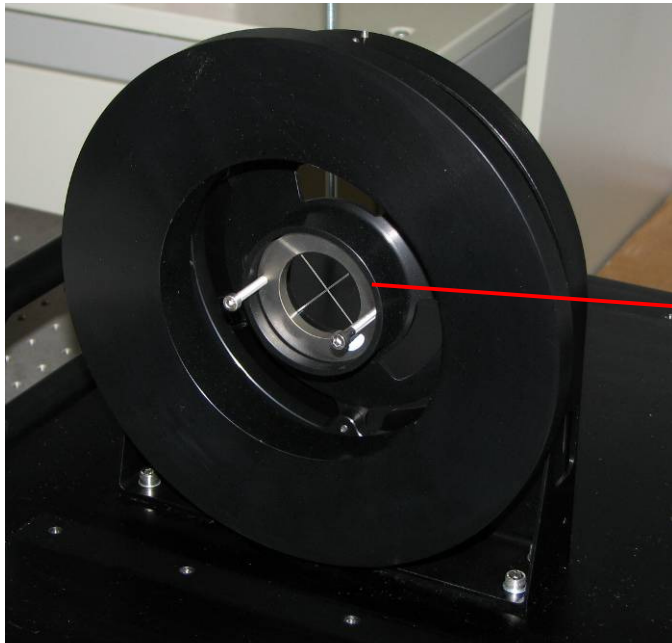
| Mode N | FREQ Hz | TRASLATIONAL | | |
|--------|---------|---------------|---------------|---------------|
| | | X | Y | Z |
| 1 | 152 | 4.64% | 1.40% | 41.27% |
| 2 | 163 | 59.20% | 0.72% | 0.71% |
| 3 | 177 | 0.64% | 33.06% | 0.00% |
| 4 | 194 | 0.65% | 0.52% | 3.17% |
| 5 | 210 | 9.44% | 0.83% | 4.36% |
| 6 | 216 | 3.31% | 0.25% | 0.72% |
| 7 | 227 | 0.00% | 0.11% | 5.82% |
| 8 | 244 | 3.57% | 4.34% | 6.36% |
| 9 | 250 | 5.85% | 5.99% | 18.29% |

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OPTICAL ALIGNMENT APPROACH



Materialization of the optical axis by means of open targets temporarily installed on the mirrors.

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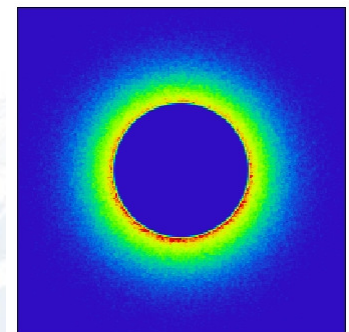
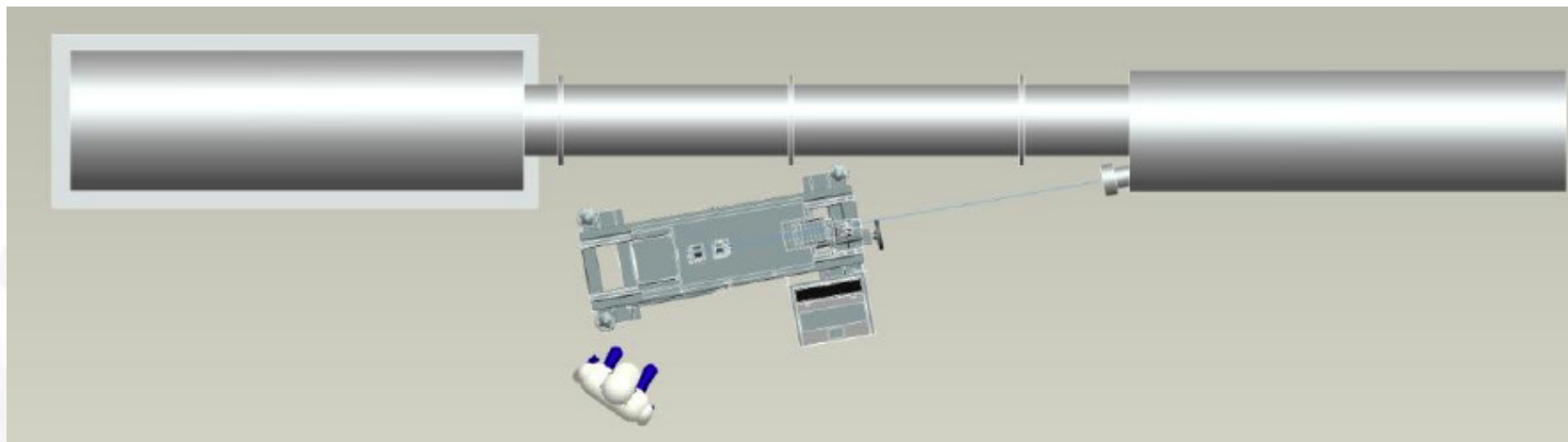
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THE ROLE OF THE OPSYS FACILITY

- Clean area (class 100) for the optical integration of METIS in controlled environment.
- Vacuum chamber for operating METIS in orbital-like environment and for hardware bake-out at controlled temperature for removing residual molecular contaminants.
- UV/EUV and VL illumination systems for the end-to-end verification of the optical alignment in all channels, for the METIS performance test (including stray-light rejection capability) and final calibration before its delivery.



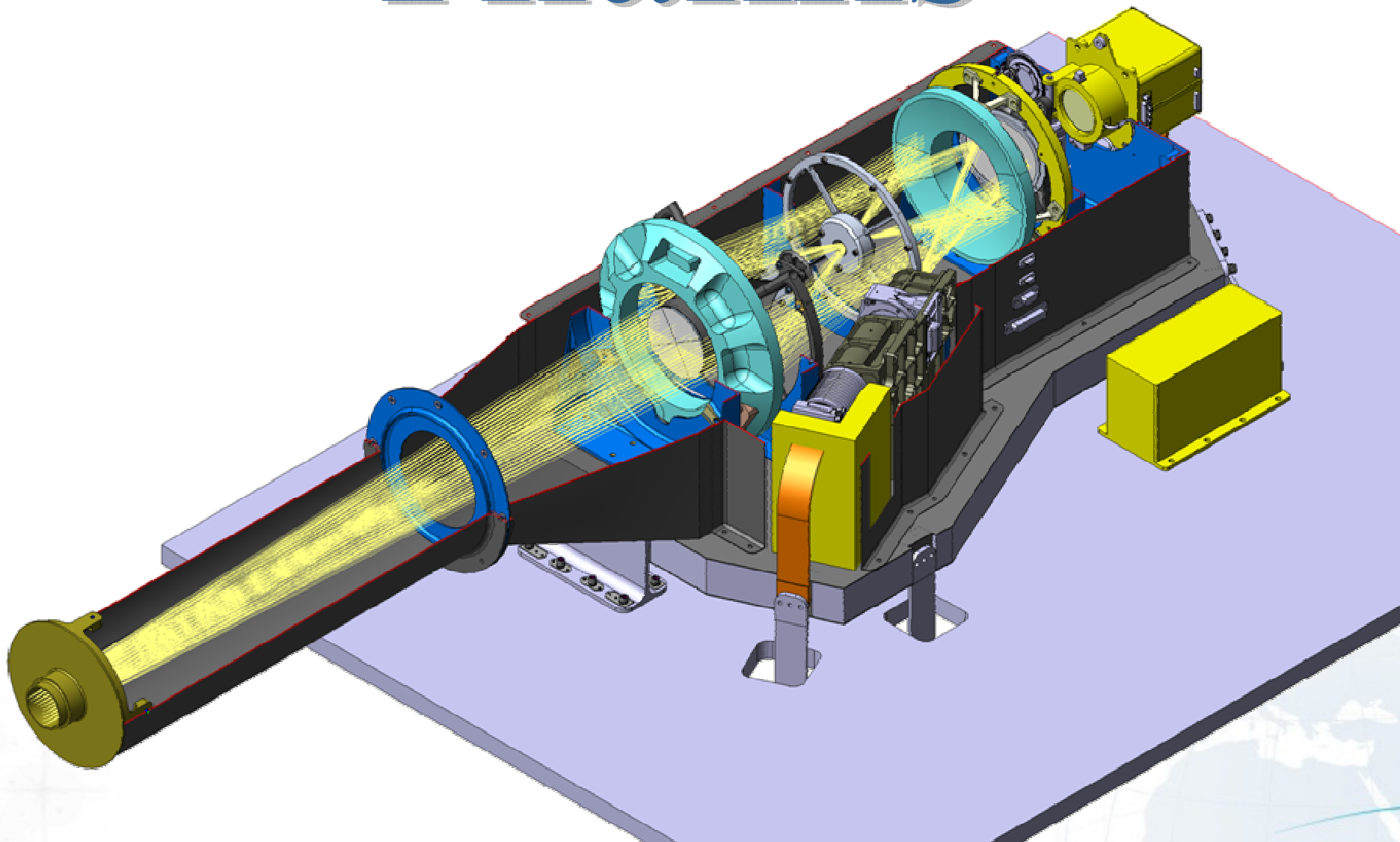
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