

# Two mirror modules

- 133 layers of grazing incidence shells with multilayer coating
- the inner 66 layers are comprised of 6 segments with 12 pieces of glass per layer
- while the outer 67 layers are comprised of 12 segments with 24 pieces of glass per layer
- Each unit is 47.2 cm) long and, upon completion, will be 19.1
- weigh 31 kg



About 5000 segments produced via hot slumping at Goddard and coated at DTU

# First light images of Cygnus X-1 on June 28th.



HEW compliant with ground calibrations (60 arcsec )

# Technology Demonstrator of SIMBOL-X/NHXM

3 mirror shells,  
600 mm length,  
10m focal length:

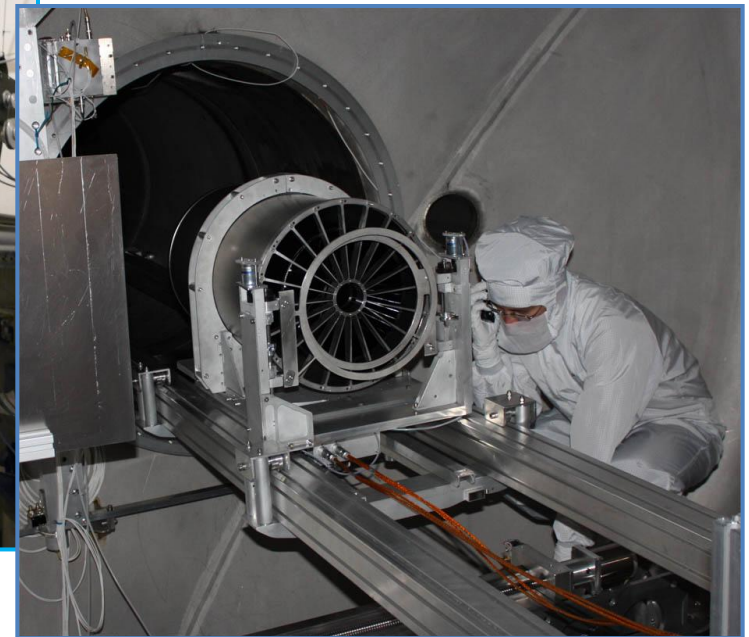
	#1	#2	#3
Diameter	286	291	297
thickness	0.25	0.25	0.25
Multi-layer	90 W-Si	200 W-Si	200 W-Si



Mechanical structure: 2  
spiders with 20 spokes



Integration of three  
mirror shell

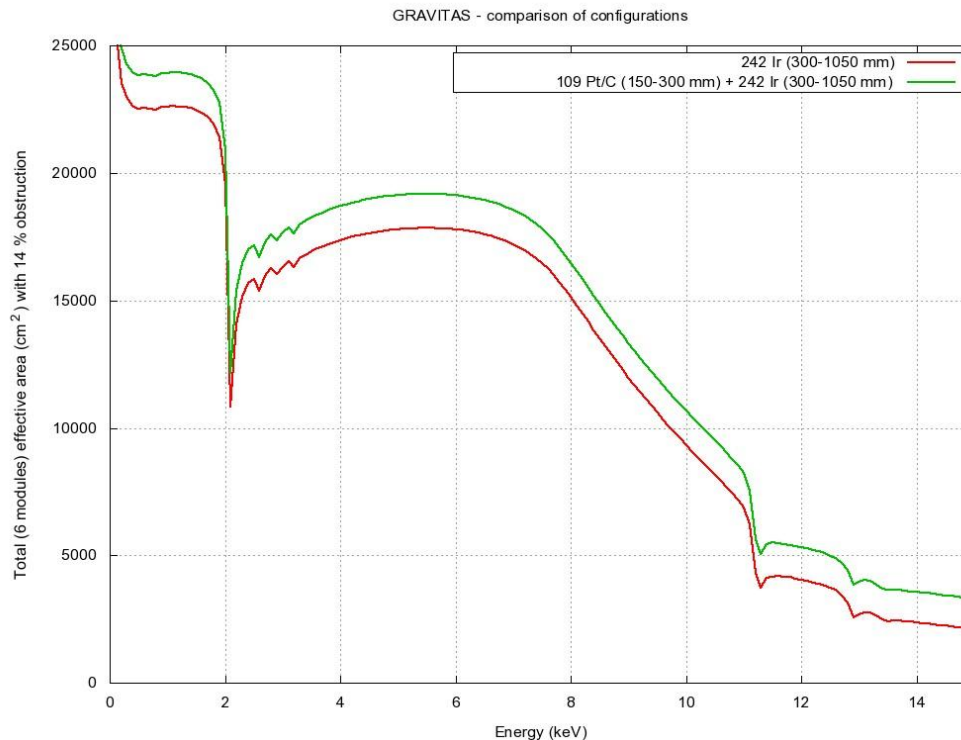


PANTER tests

# NHXG (NEW HARD X-RAY GRAVITAS)

1 single module = 1 NHXM + 1  
GRAVITAS units (nested together)  
Mirror material: Ni/glass segments

MASS (1 MODULE)  $\approx$  200 kg

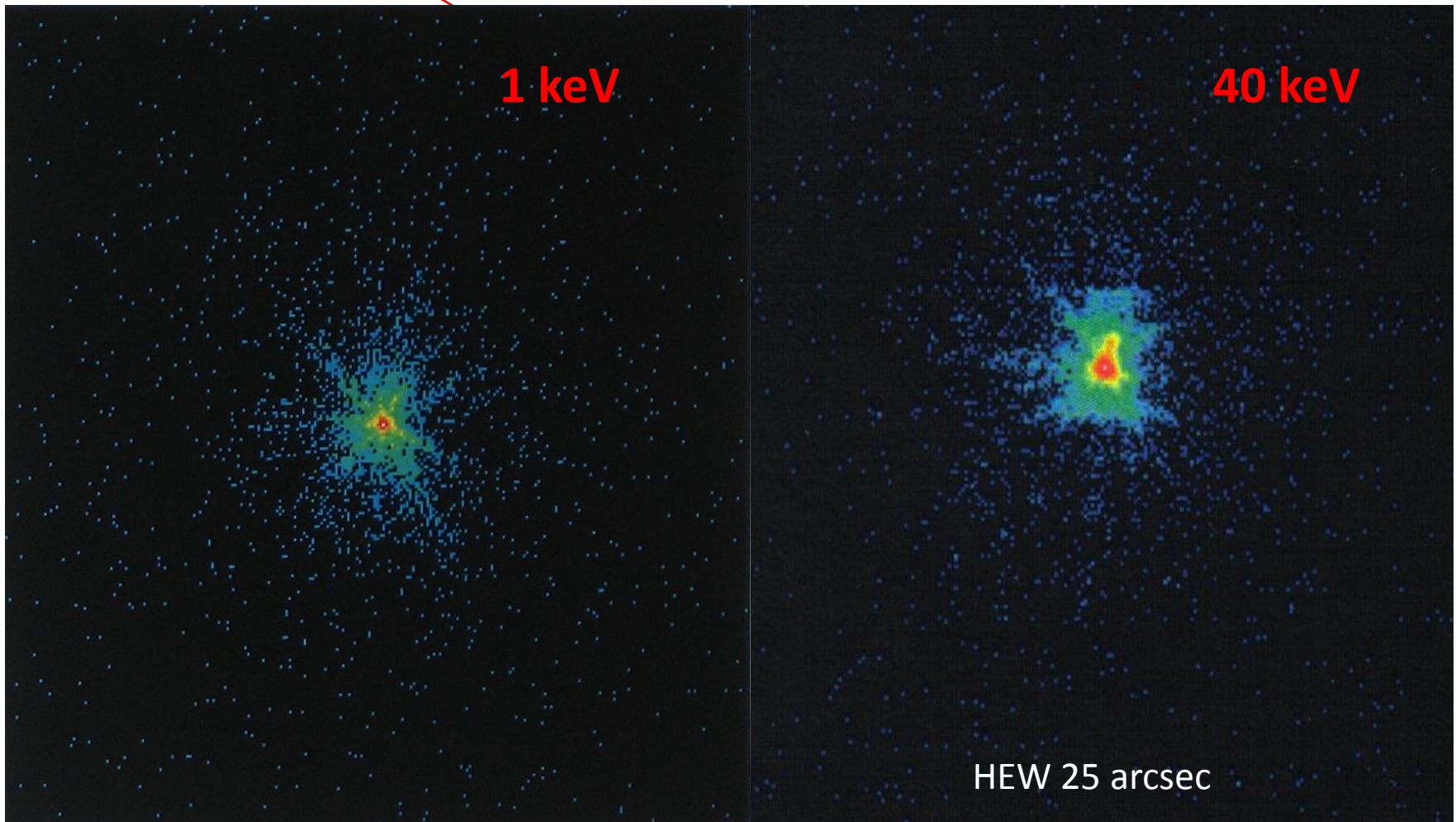


6 mirror modules  
HEW  $\rightarrow$  10-20 arcsec

NB: Possible improvement with polarimetry and extension of the band with Bragg crystals



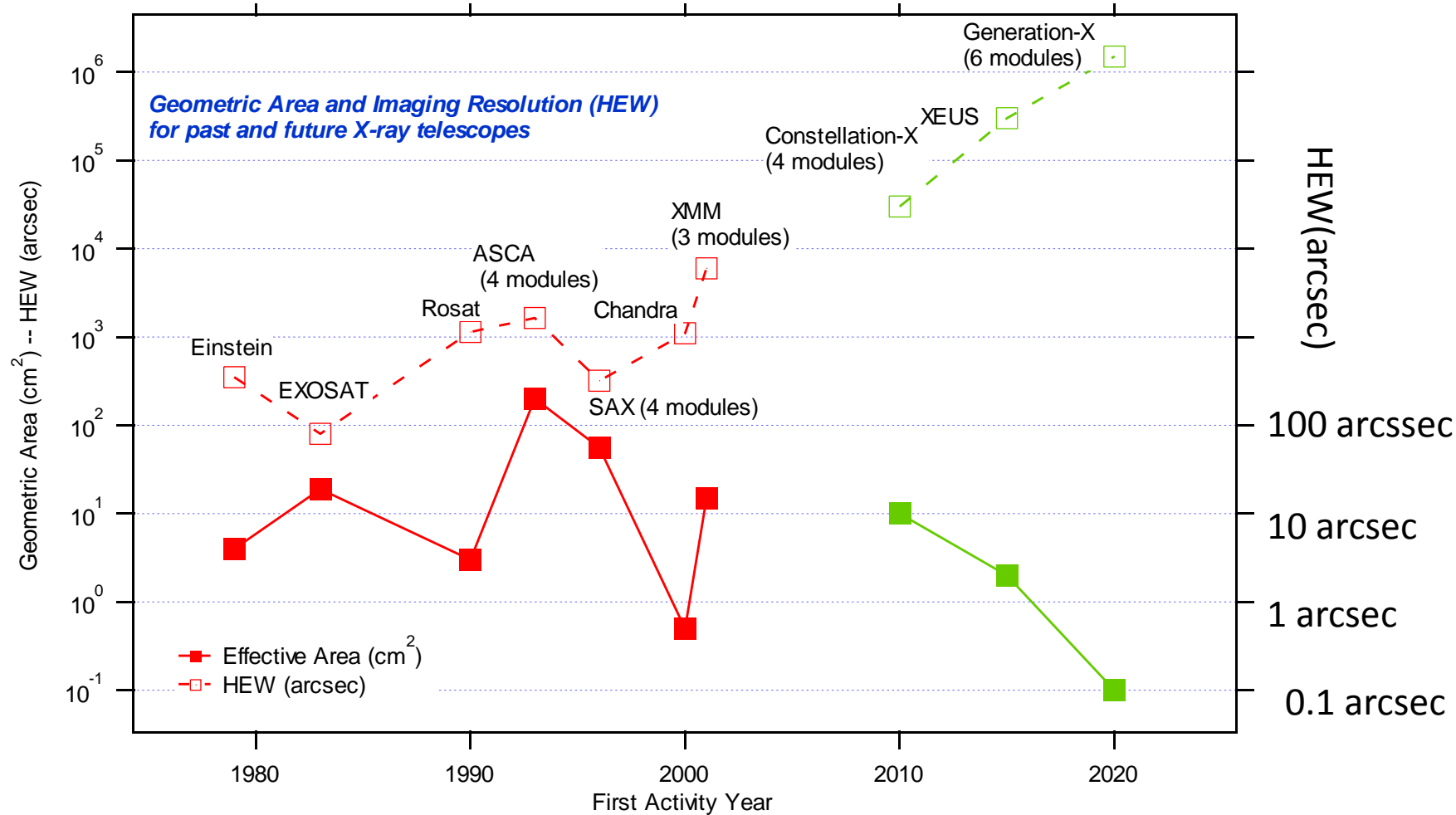
# Optics calibration @ Panter/MPE



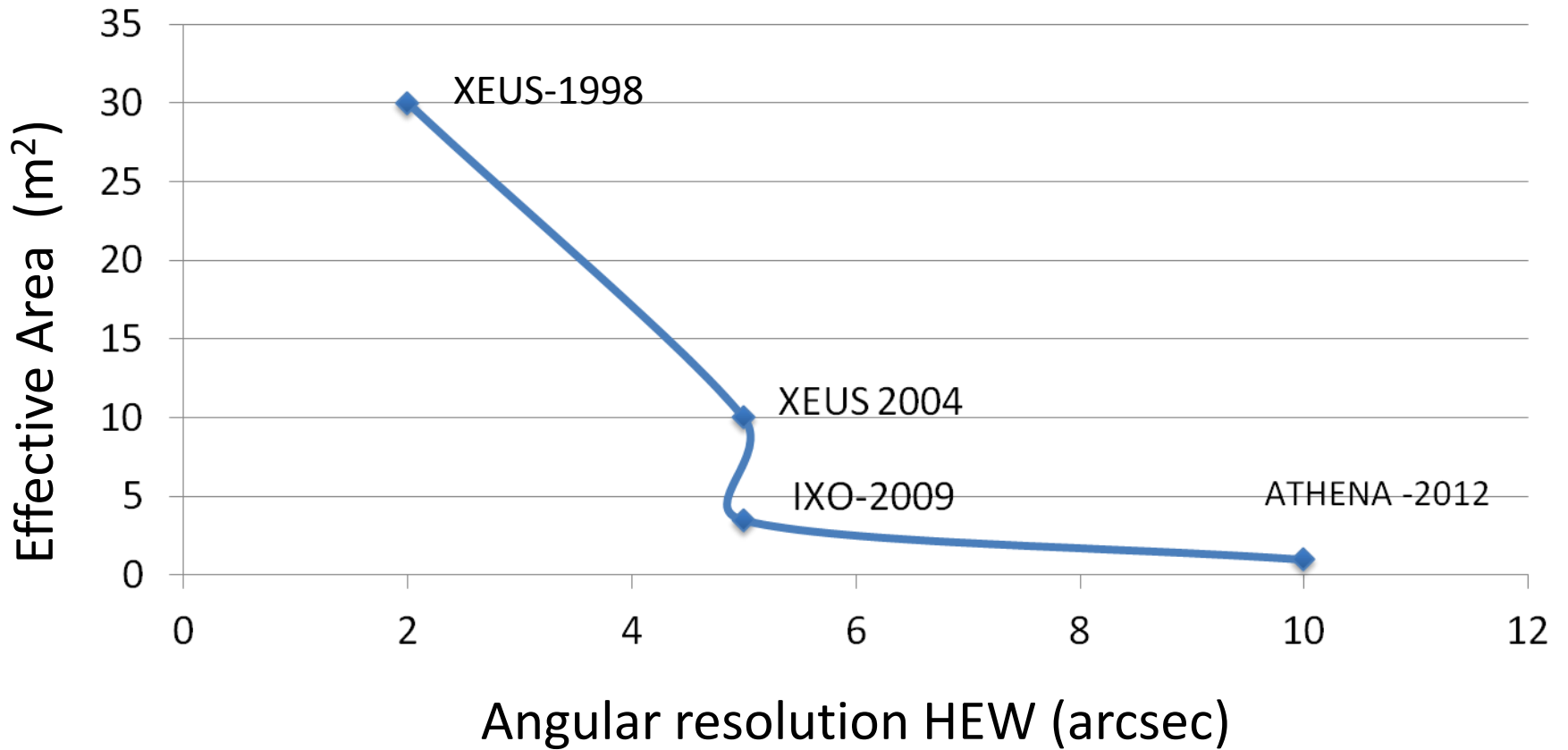
# Angular resolution for Hard X-ray Experiments & Telescopes

Experiment - Telescope	Year	“Imaging” technique	Angular resolution
SAX-PDS	1996	Rocking collimator	> 3600 arcsec (collimator pitch)
INTEGRAL-IBIS	2002	Coded mask	720 arcsec (mask pitch)
<b>NUSTAR</b>	<b>2012</b>	<b>Multilayer Wolter I mirrors</b>	<b>60 arcsec HEW</b>
<del>NHXM/SIMBOL -X-like?</del>	?	Multilayer Wolter I Optics	> / >> 20 arcsec HEW

# Geometric Area and Angular resolution for past and future X-ray telescopes



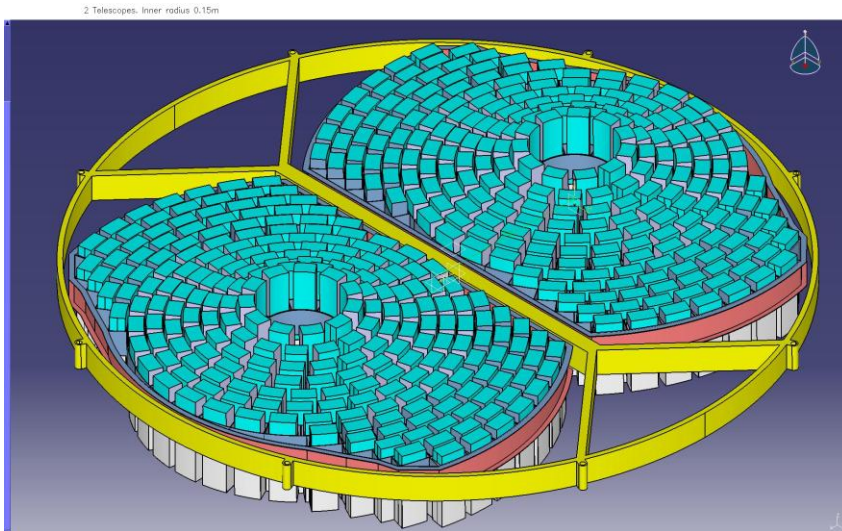
## XEUS/IXO/ATHENA Effective Area Vs. HEW



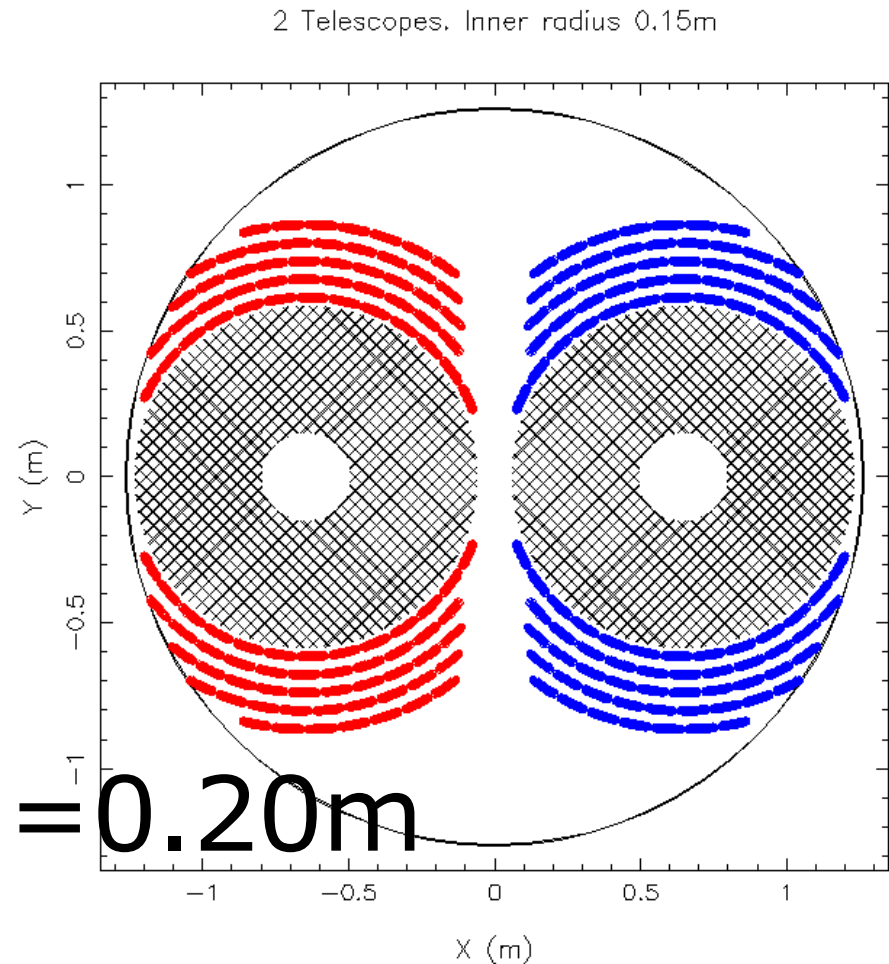
Is the science case changed across the time?



# ATHENA - "Owl Eyes" Design



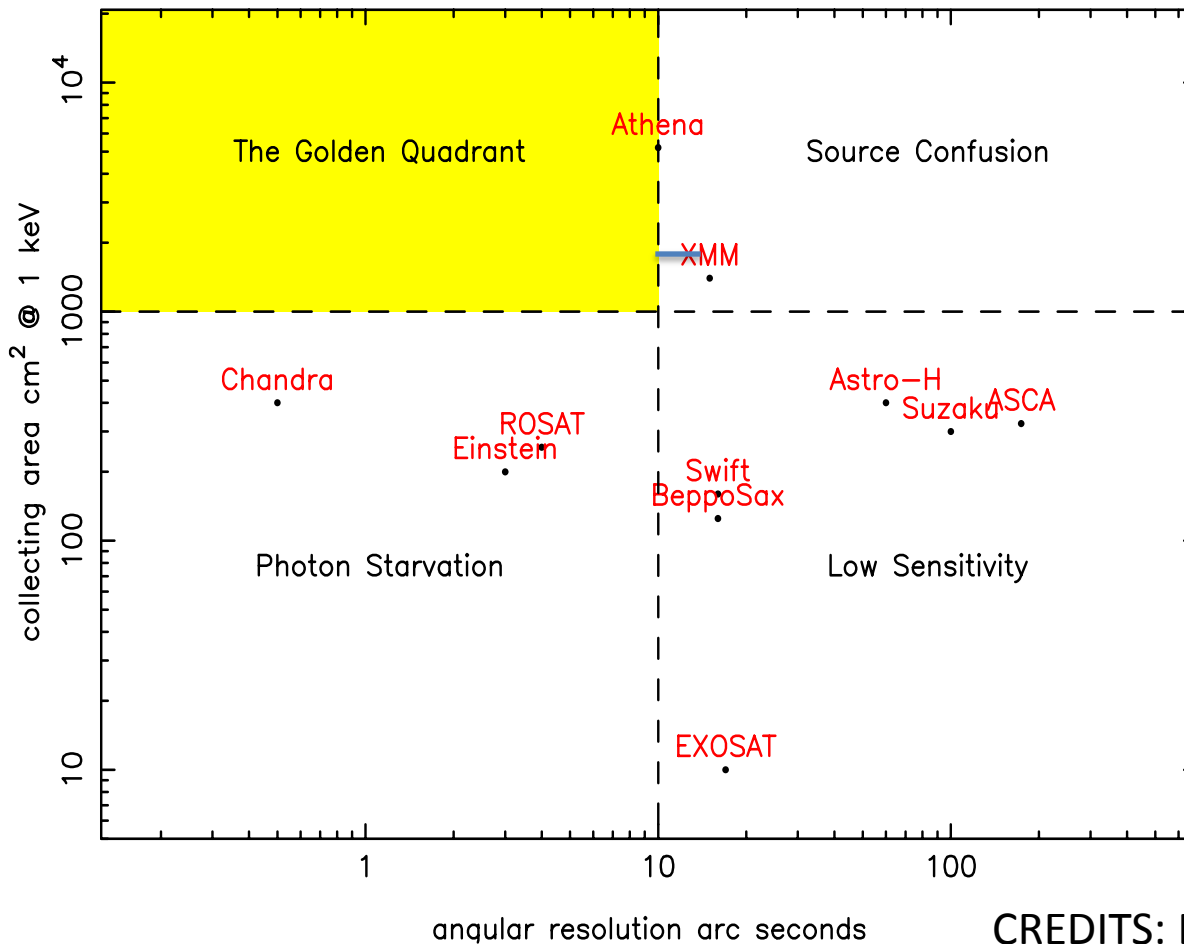
HEW requirement = 10 arcsec



**$F=11.5\text{m}$  and  $R_{in}=0.20\text{m}$**

# Athena Telescope Area and Resolution

The performance of X-ray Telescope Modules



CREDITS: Dick Willingale



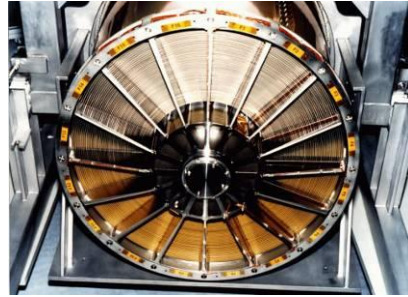
CHANDRA

0.5 arcsec

$A_{\text{eff}} = 0.078 \text{ m}^2$

$18500 \text{ kg/m}^2$

$A_{\text{eff}} @ 1 \text{ keV}$



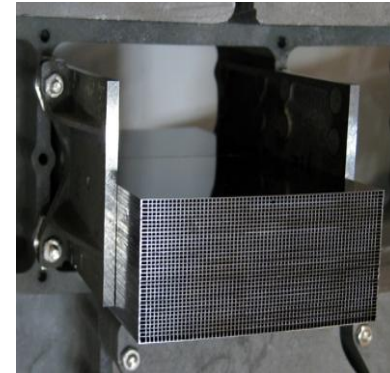
XMM-NEWTON

10 arcsec

$A_{\text{eff}} = 0.43 \text{ m}^2$

$2300 \text{ kg/m}^2$

$A_{\text{eff}} @ 1 \text{ keV}$



Silicon Pore Optics

10 (5) arcsec

$A_{\text{eff}} = 1 \text{ m}^2$

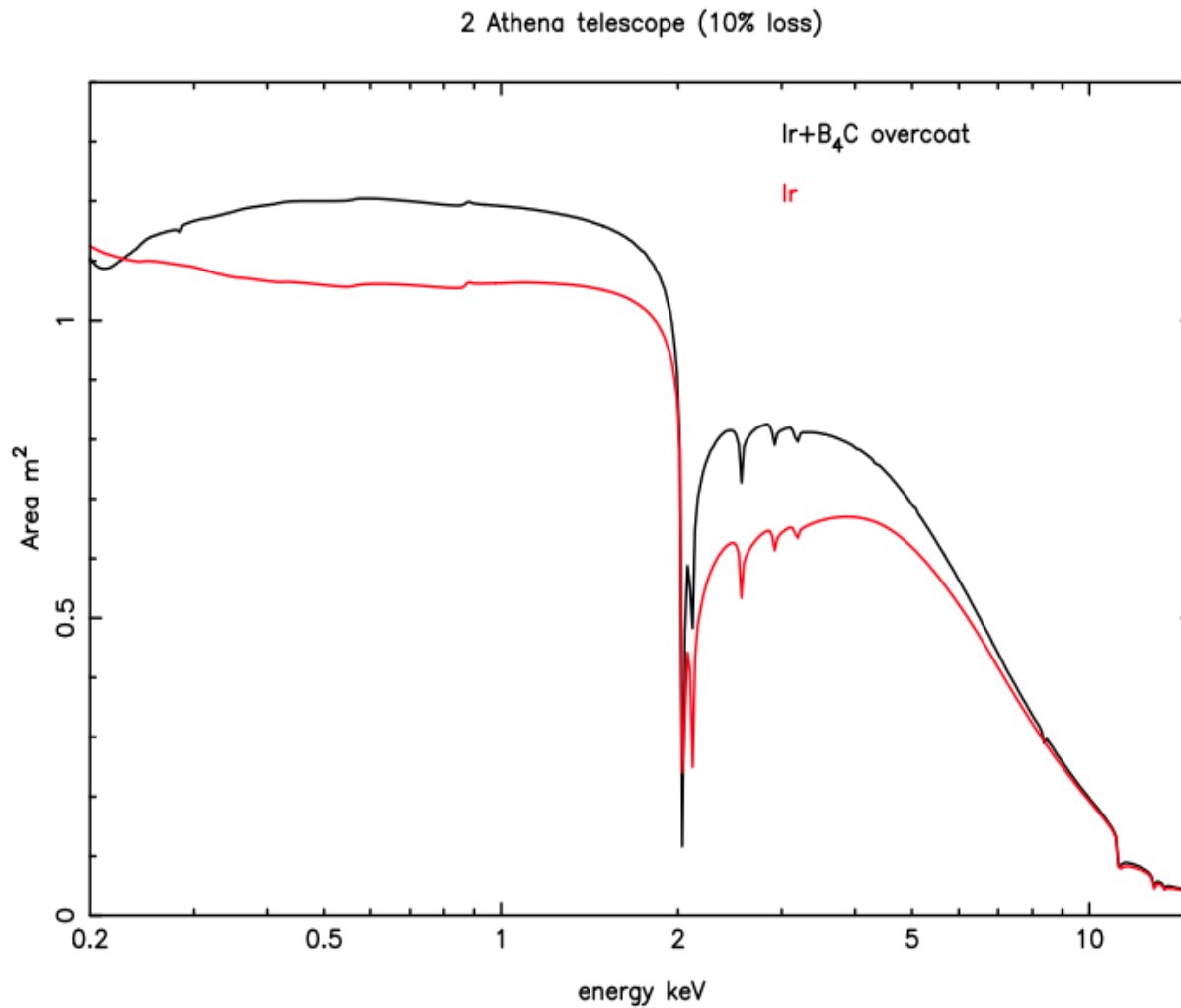
$< 240 \text{ kg/m}^2$  (mirror module)

$2500 \text{ kg/m}^2$  mirror system

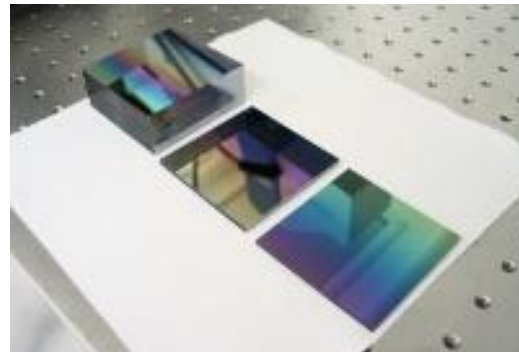
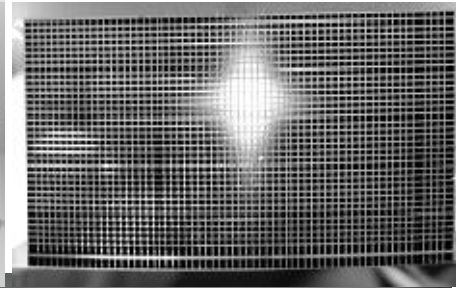
$A_{\text{eff}} @ 1 \text{ keV}$

**ATHENA baseline**

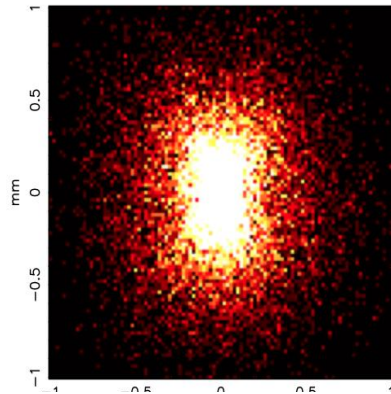
# ATHENA's effective area



# PORE OPTICS TECHNOLOGY

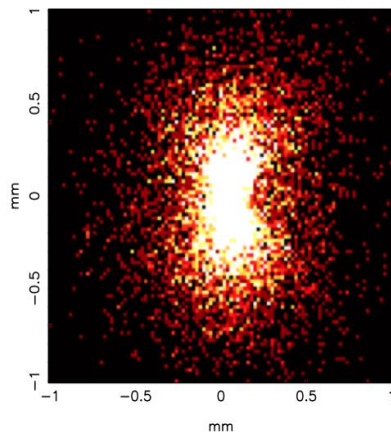


# Si-Pore Optics PSF



PSF on-axis at 1.25 keV. The rms major/minor ratio is 1.21

Asymmetry introduced because the aperture is not circular



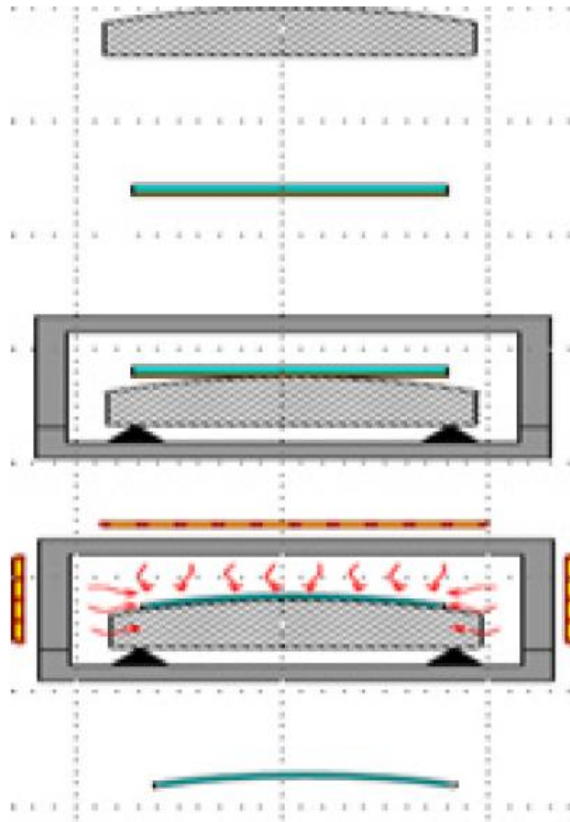
PSF 10 arc minutes off-axis at 1.25 keV. The rms major/minor ratio is 1.45

Current calibrations at Panter: 20-25 arcsec HEW for the single unit (about 20 cm<sup>2</sup>)  
Improvements aiming at reaching 5 – 10 arcsec

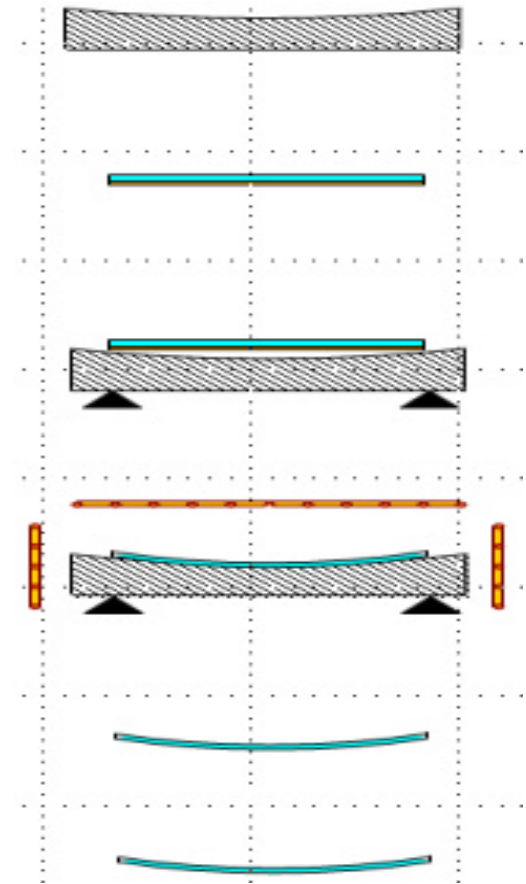


# Direct & Indirect slumping processes

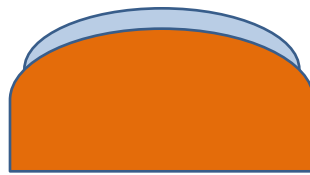
DIRECT



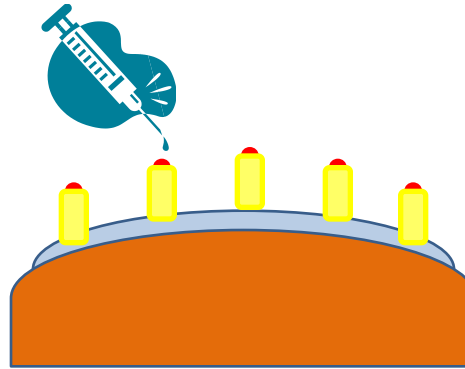
INDIRECT



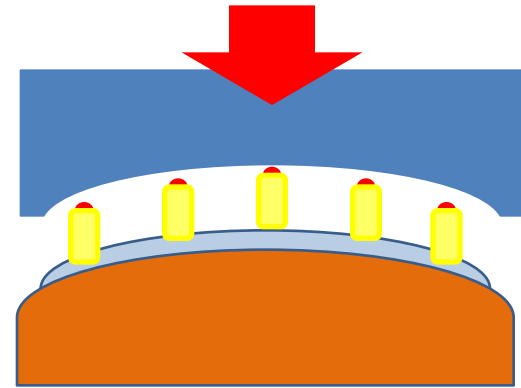
# Slumped glass integration



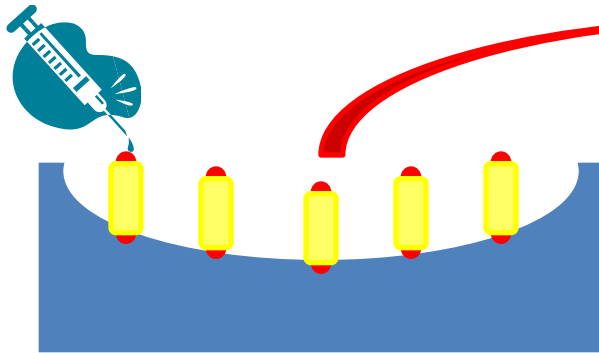
Glass fixed  
To the mould



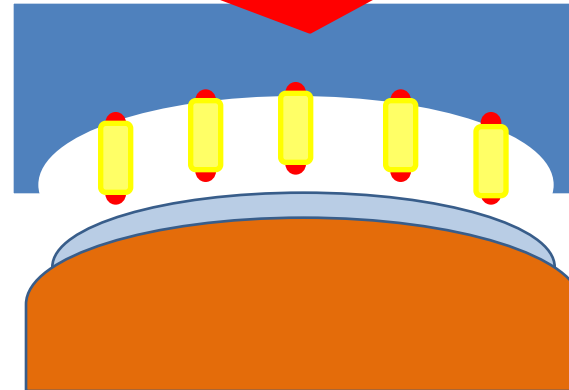
Ribs aligned on the glass  
And glue dispensed



Ribs glued to backplane



Glue dispensed on the ribs



Backplane with  
ribs glued to  
glass fixed on  
mould

# Slumping Glass results (OAB/ESA)

## 2.1.1. Characterization before edge trimming

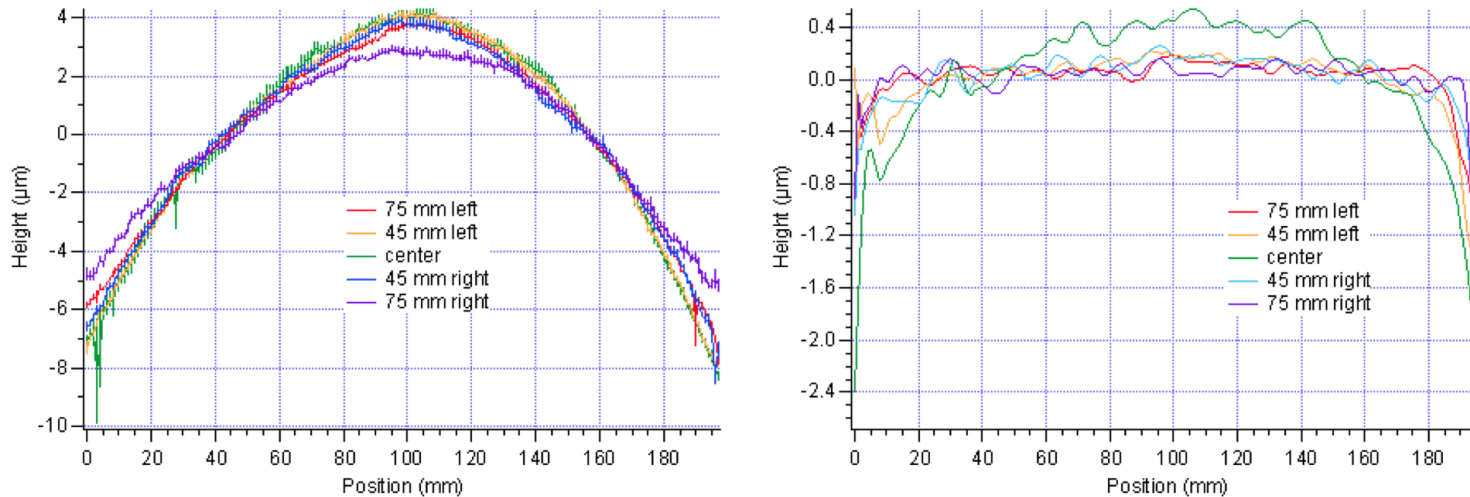
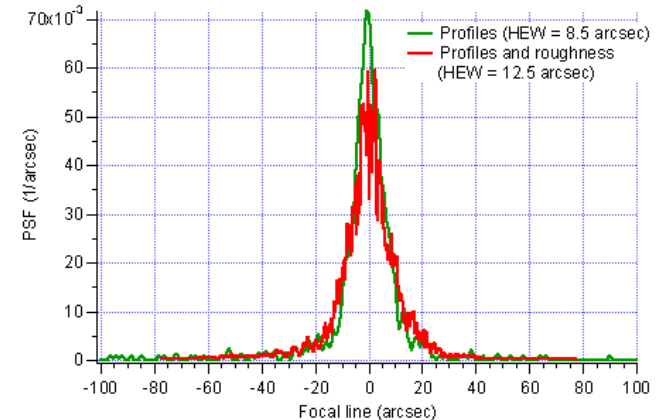
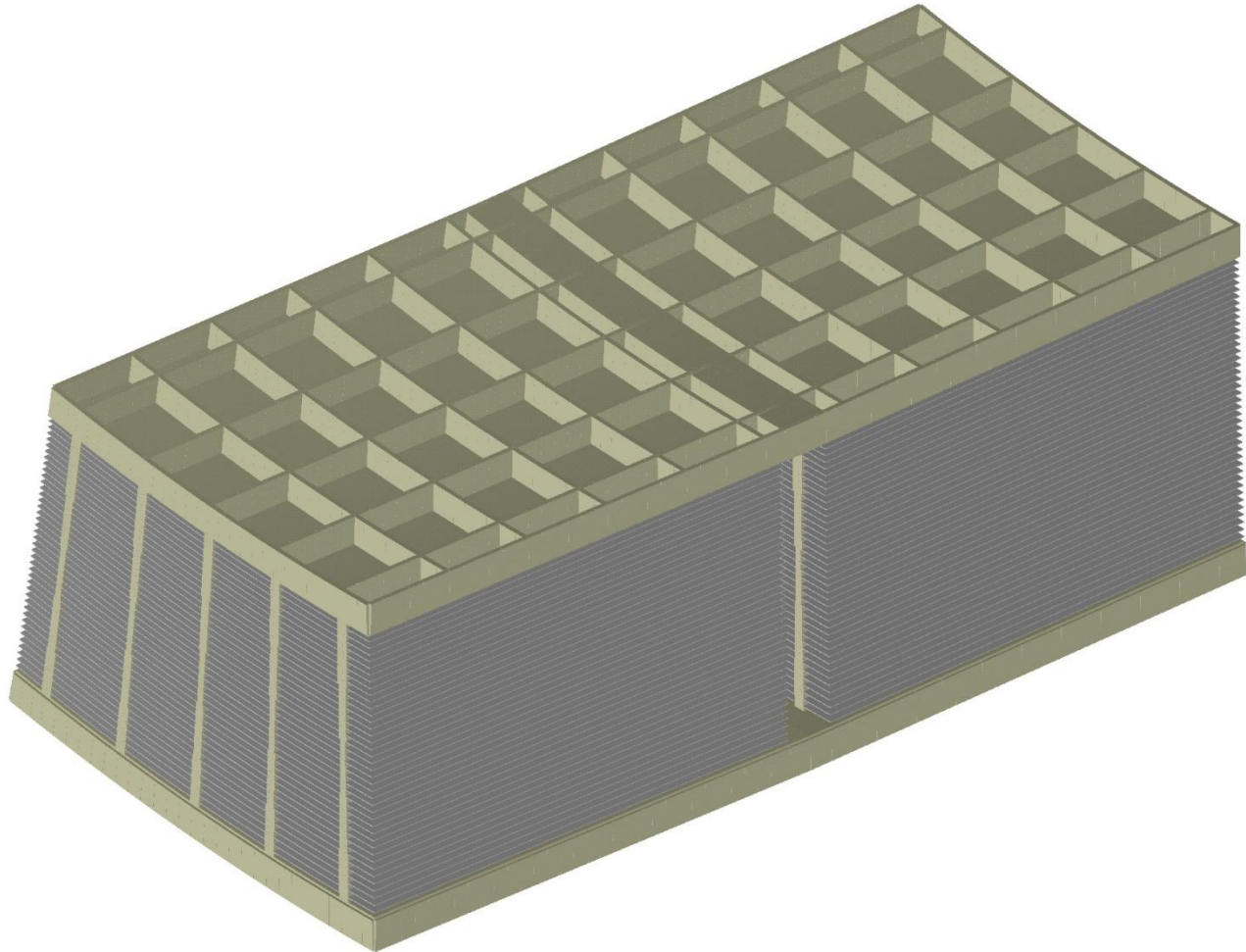


Fig. 1: (left) measured longitudinal of the PS-IXO-196 with the LTP+CHR, over a 200 mm scan length. (right) after simulating the integration effect.

HEW  $\approx$  10 arcsec

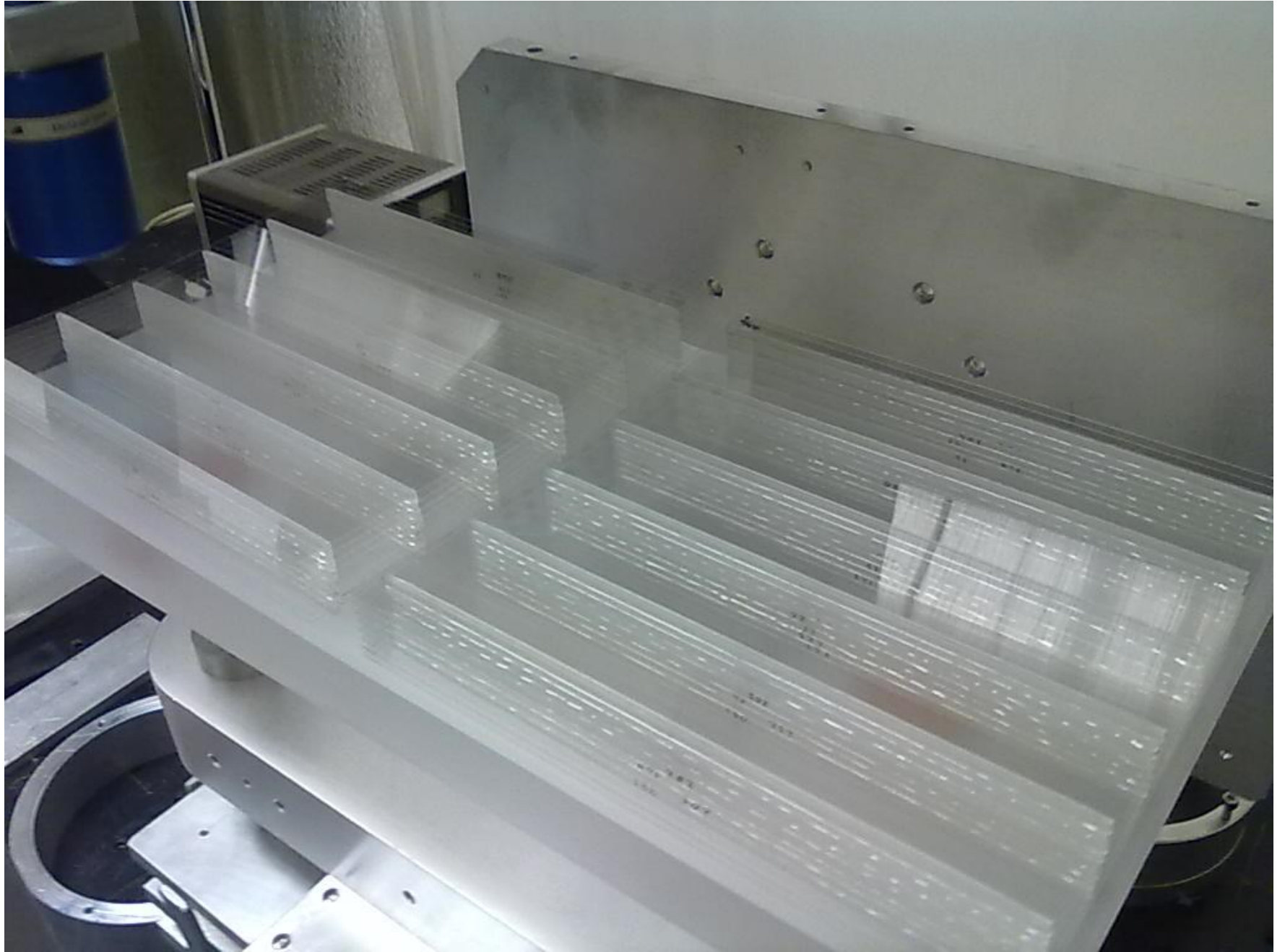


# THE INTEGRATION CONCEPT



IXO - STACK

**XOU-BB**



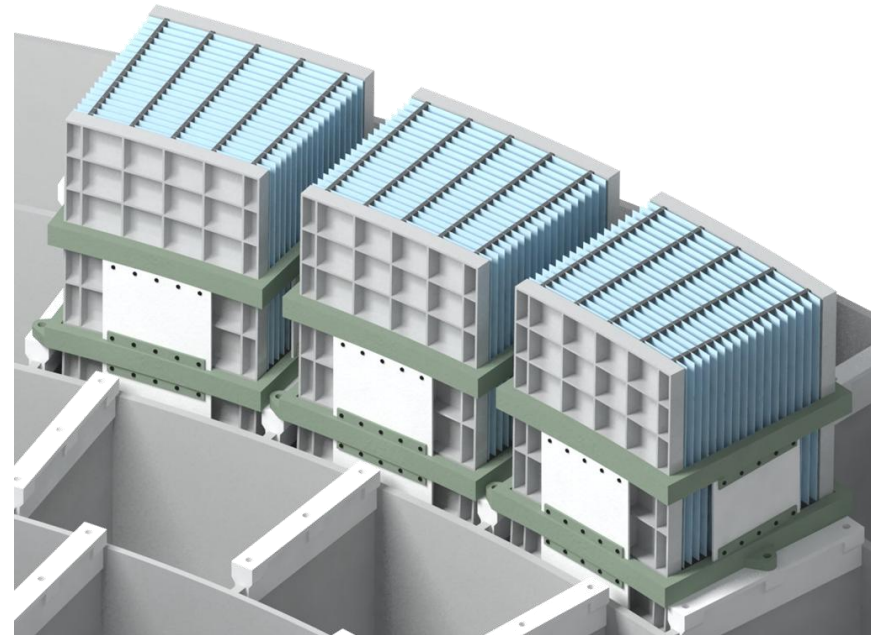
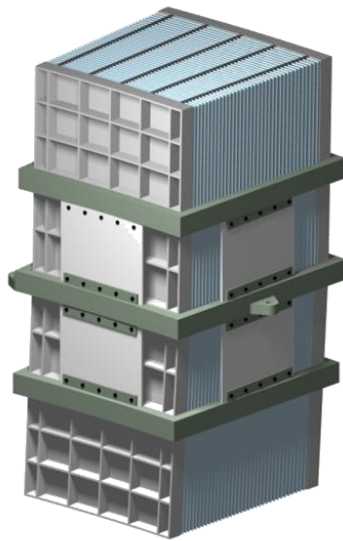
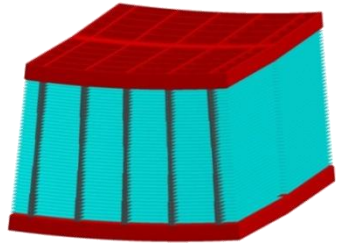


## 8 plate pairs (PP1, PP2... PP8)





From the double plated stack (with ribbed glass back-plates) to the optics module assembly



# Flight Mirror Assembly

Hierarchy principle for fabrication of the complete mirror assembly

## Main Parameters for the Telescope Configuration

Number of MS per XOU	35-60
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Number of Rings	9
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Number of XOUs	246
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MS thickness	0.4mm
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Ribs average thickness	4.2mm
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Average XOU mass	~ 10.5 Kg
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